Discussion Session @ MIAPP, 26 October 2016

## Rare B Decays: Standard Model vs. New Physics

David M. Straub Universe Cluster/TUM, Munich



## Status of $b \rightarrow s\mu^+\mu^$ anomalies

David Straub (Universe Cluster)

・ロト (四) (日) (日) (日)

### Current tensions in $b \rightarrow s\mu^+\mu^-$ transitions

Mode	Observable	Bin	Pull
$B^0  ightarrow K^* \mu^+ \mu^-$	$P_5'$	4-6	<b>-2.6σ</b>
$B_{ m s}  o arphi \mu^+ \mu^-$	BR	1-6	<b>-3.3σ</b>
$B^+  ightarrow K^+ \mu^+ \mu^-$	BR	1-6	<b>-2.0σ</b>
$B^+  ightarrow K^+ \mu^+ \mu^-$	BR	15-22	<b>-2.6σ</b>

#### Suspects: New physics? Form factors? Charm loop?

(flavio v0.13.1 using combined LCSR+LQCD FFs for  $B \rightarrow V$  FFs Bharucha et al. 1503.05534 and FNAL/MILC  $B \rightarrow K$  FFs Bailey 1509.06235; hadronic unc. estimated as in Altmannshofer and Straub 1411.3161)

David Straub (Universe Cluster) < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □

# New physics: effective field theory

David Straub (Universe Cluster)

▲□▶▲圖▶▲≣▶▲≣▶ = 亘 - のへ⊙

### Global constraints on C<sub>9</sub> & C<sub>10</sub>

- including 3 fb<sup>-1</sup> LHCb measurements
- ► Updated B → V FFs from v2 of Bharucha et al. 1503.05534
- Best fit point: 4.5σ pull from SM

What does it mean?



see also Altmannshofer and Straub 1411.3161, Descotes-Genon et al. 1510.04239, Hurth et al. 1603.00865

David Straub (Universe Cluster)

▲ロト ▲屈 ト ▲ 臣 ト ▲ 臣 ト ● ④ ● ● ●

### A closer look

Pulls for individual modes:

- ►  $B \rightarrow K^* \mu^+ \mu^-$ : 2.7 $\sigma$ 
  - ▶ famous P'<sub>5</sub> anomaly
- ►  $B_s \rightarrow \varphi \mu^+ \mu^-$ : **3.4** $\sigma$ 
  - BR @ low & high q<sup>2</sup>
     cf. Bharucha et al. 1503.05534,
     Ronald R. Horgan et al. 1310.3887
- ►  $B \rightarrow K\mu^+\mu^-$ : **2**.6 $\sigma$ 
  - ▶ BR @ low  $q^2 \rightarrow R_K!$
  - First pointed out in:

Khodjamirian et al. 1211.0234



### Key question

Given new physics can account for the tensions, are we sure there is no SM effect that can explain them?

 $\equiv$  Do we have our theory uncertainties under control?

# Scrutinizing theory uncertainties: form factors

David Straub (Universe Cluster)

▲ロト ▲屈 ト ▲ 臣 ト ▲ 臣 ト ● ○ ● ● ● ●

### $B \rightarrow K^*$ form factors

#### Complementarity and agreement between LCSR & LQCD form

factors Bharucha et al. 1503.05534, R. R. Horgan et al. 1501.00367



David Straub (Universe Cluster)

### $B \rightarrow K^*$ form factors

#### Complementarity and agreement between LCSR & LQCD form

factors Bharucha et al. 1503.05534, R. R. Horgan et al. 1501.00367



David Straub (Universe Cluster)

### $B \rightarrow K^*$ form factors

## Complementarity and agreement between LCSR & LQCD form factors Bharucha et al. 1503.05534, R. R. Horgan et al. 1501.00367



David Straub (Universe Cluster)

### **Discussion points**

- Can we trust LCSR FFs to 10% (and their ratios to 5%)?
- ► Issue of K\* instability in LQCD and LCSR?
- ▶ LCSR with K\* LCDAs vs. B LCDAs

# Scrutinizing theory uncertainties: non-factorisable effects

David Straub (Universe Cluster)

▲□▶ ▲□▶ ▲□▶ ▲□▶ = 三 - つく()

### Non-factorisable effects: "charm loop"

Culprit: matrix element of O<sub>1,2</sub>

 $\langle \bar{K}^* | T\{j^{\mu}_{em}(x)C_{1,2}O_{1,2}(0)\} | \bar{B} \rangle$ 

- Since O<sub>9</sub> ∝ ℓ̄γ<sup>μ</sup>ℓ, h<sub>λ</sub> could mimic a new phyiscs effect in C<sub>9</sub>
- Without loss of generality, can be parametrised as complex-valued (CP-even) helicity-dependent shift in C<sub>9</sub>:

$$C_9^{\rm SM} + \Delta C_9^{+,-,0}(q^2)$$

$$O_2 = (\bar{s}_L \gamma_\mu c_L)(\bar{c}_L \gamma^\mu b_L)$$







DS @ Moriond EW 2015; Altmannshofer and Straub 1503.06199 ( $1\sigma$  boxes)

David Straub (Universe Cluster)

 $q^2$  dependence of  $\Delta C_{\rm o}^{\lambda}$ 



Ciuchini et al. 1512.07157

- Bayesian fit assuming a polynomial form for  $h_{\lambda}$
- roughly:  $\tilde{g}_1 \propto \Delta C_9^-$ ,  $\tilde{g}_3 \propto \Delta C_9^0$

David Straub (Universe Cluster)

### $q^2$ dependence of $\Delta C_9^{\lambda}$



#### Ciuchini et al. 1512.07157

- Bayesian fit assuming a polynomial form for  $h_{\lambda}$
- roughly:  $\tilde{g}_1 \propto \Delta C_9^-$ ,  $\tilde{g}_3 \propto \Delta C_9^0$
- **assuming** small  $\Delta C_9^{\lambda}$  for small  $q^2$  (expected for SM, but not NP!)

 $q^2$  dependence of  $\Delta C_{\rm o}^{\lambda}$ 

- Bin-by-bin fit of  $\Delta C_9^0$  vs.  $\Delta C_9^-$
- New physics: expect ΔC<sub>9</sub><sup>0</sup> = ΔC<sub>9</sub><sup>−</sup> equal for all bins

Current data **not precise enough** to exclude new physics hypothesis!



Plot based on discussion with C. Bobeth

### **Dispersion relation**

 Charm contribution obeys a dispersion relation Khodjamirian et al. 1006.4945

Schematically:

$$h_{-}(q^{2}) = h_{-}(0) + q^{2}h'_{-}(0) + q^{4} \left[ \underbrace{\mathsf{BW}_{J/\psi} + \mathsf{BW}_{\psi(2S)}}_{\substack{\mathsf{Measured from } B \to \psi K^{*} \\ \mathsf{up to overall phase}}} + \underbrace{h_{-}^{\mathsf{higher}}(q^{2})}_{\substack{\mathsf{small impact} \\ \mathsf{below } m_{J/\psi}^{2}(?)}} \right]$$

etc.

David Straub (Universe Cluster)

### Charmonium interference

The  $q^2$  dependence of the differential rate between the  $J/\psi$  and  $\psi(2S)$  resonances can be used to infer the sign of the interference

Khodjamirian et al. 1006.4945, Lyon and Zwicky 1406.0566



\* this is only a cartoon - not actual numerics

### Meanwhile at LHCb

#### Measuring phase differences

- Measure relative phase between narrow resonances and penguin amplitudes
   Model resonances as relativistic BWs multiplied by relative scale and phase Lyon et al. [1406.0566], Hiller et al. [1606.00775]
  - ightarrow Use this model to replace  $Y(q^2)$  in  $C_9^{eff}=Y(q^2)+C_9$
  - $\rightarrow$  B  $\rightarrow$  K form factors constrained to LCSR+Lattice predictions
  - $\rightarrow$  Fit for phases and  $\mathit{C}_{9}$  and  $\mathit{C}_{10}$







### **Discussion points**

- Exploiting the dispersion relation (and B → ψK\* data), can we put an upper (lower) bound on the charm loop effect in the critical 4−6 GeV<sup>2</sup> region?
- Any other way to get a handle on the size (and sign) of the effect?

### Outlook

- Which measurements at Belle-II could help reducing theory uncertainties?
- Which measurements at Belle-II could help probing a possible NP effect?
- ▶ Which cleaner observables will be accessible at Belle-II?
  - inclusive  $B \to X_{s} \ell^{+} \ell^{-}$
  - ▶ ...