#### High-speed electroweak baryogenesis



Jim Cline, McGill University based on JC, K. Kainulainen 2001.00568 webinar, Stefan Institute, Ljubljana, 26 May, 2020

# **Electroweak baryogenesis**

EWBG relies on a strongly 1st order electroweak phase transition, and CP-violating interactions of fermions at the bubble walls,



Needs new physics at the electroweak scale to get both ingredients.

Recently high wall velocities  $v_w$  became more interesting because of gravity waves. Can EWBG work at high  $v_w$ ?

#### Lore

It was believed that EWBG gets quenched if  $v_w \rightarrow 1/\sqrt{3}$ , the speed of sound.

PHYSICAL REVIEW D

VOLUME 53, NUMBER 6

15 MARCH 1996

Nonlocal electroweak baryogenesis. II. The classical regime

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The fluid equations (26) are calculated to leading order in  $v_w$ . A fuller analysis incorporating all orders in  $v_w$ can be performed and shows that the velocity at which the leading-order analysis breaks down is the speed of sound  $v_w \sim v_s = 1/\sqrt{3}$  in the plasma. If the wall moves faster than this, there is no solution in front of the wall, and perturbations cannot propagate into this region. We are interested in the case when perturbations can propagate in front of the wall where the anomalous electroweak processes are unsuppressed. Thus, we assume

$$v_w < v_s = \frac{1}{\sqrt{3}} \,. \tag{31}$$

#### Lore

# It was believed that EWBG gets quenched if $v_w \rightarrow 1/\sqrt{3}$ , the speed of sound.

PHYSICAL REVIEW D

VOLUME 53, NUMBER 8

15 APRIL 1996

#### Electroweak baryogenesis in supersymmetric models

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tion. Should the wall velocity approach or be larger than the speed of sound, diffusion is not a good approximation to transport and our computations are invalid. An improved cal-

People thought that particles in front of a supersonic wall don't have time to diffuse away from the wall.

But is it true? Sound propagation and diffusion are entirely different phenomena!

#### **Simple exercise**

Compute the fraction F of particles in a thermal gas that have  $v_z = p_z/E > v$  (can stay ahead of the wall) as a function of v.

$$F = \frac{\frac{1}{2} \int_{\gamma_w v_w m}^{\infty} dp \, p^2 (1 - v_w E/p) / (e^{\beta E} + 1)}{\int_0^\infty dp \, p^2 / (e^{\beta E} + 1)}$$



Nothing special happens at  $v = 1/\sqrt{3}!$  Diffusion only needs random particle motions.

#### **Our result**

A consistent derivation of fluid equations leads to physically sensible result: diffusion can occur up to  $v_w \leq 1$ .

If your fluid equations predict something else, there is a problem.

Hence EWBG can in principle work for  $v_w > 1/\sqrt{3}$ 

Same phase transition might produce both EWBG and gravity waves observable by LISA

But first some more general background on EWBG ...

# How to get a strong phase transition?

First order phase transition requires potential barrier,



Traditionally, the barrier came from finite-temperature cubic correction to potential,

$$\Delta V = -\frac{T}{12\pi} \sum_{i} (m_i^2(h))^{3/2} = -\frac{T}{12\pi} \sum_{i} (m_{i,0}^2 + g_i^2 h^2 + c_i T^2)^{3/2}$$

It is typically not very cubic, and not big enough. Tends to give only a 2nd order or weak 1st order phase transition, v/T < 1.

# **Previous attempts**

In the past, much attention was given to EWBG in the MSSM with light stops, and two Higgs doublet models (2HDMs). These are largely defunct.



artwork credit: K. Kainulainen

No light stops observed, and EDM constraints on needed CP violation are severe.

Recent studies of 2HDM viability (Dorsch, Huber, Konstandin, No, arXiv:1611.05874) are trumped by improved EDM constraints.

#### **Demise of 2HDM EWBG** ACME (Nature 562 (2018) 355) killed 2HDM EWBG:



### **Tree-level barrier with a singlet scalar**

A more robust scenario introduces a scalar singlet *s*. Choi & Volkas, hep-ph/9308234; Espinosa, Konstandin, Riva, 1107.5441



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# **Singlet can help with CP violation**

Can introduce dimension-5 coupling to top quark,  $i(s/\Lambda)\bar{Q}_LHt_R$ , to give complex mass in the bubble wall,

$$m_t(z) = \frac{y_t}{\sqrt{2}} h(z) \left( 1 + i \frac{s(z)}{\Lambda} \right) \equiv |m_t(z)| e^{i\theta(z)}$$

Can be derived from model with heavy vector-like top partner coupling to *s*.

This gives the CP-violating interactions of t in the wall, producing CP asymmetry between  $t_L$  and  $t_R$ .

The new CP violation is spontaneous, due to  $\langle s \rangle \neq 0$  during EWPT, and disappears at  $T \leq m_W$ .

One escapes stringent constraints on new CP-violation from EDM searches.

# **Heavy top partners**

A simple UV completion is a vector-like top partner  $T_{R,L}$  coupling to singlet,

$$\eta \, \bar{t}_R S T_L + M_T \bar{T}_L T_R + y' \bar{T}_R H t_L$$

Integrate out heavy state:



Generates desired coupling

$$\frac{\eta y'}{M_T} \, \bar{t}_R S H t_L$$

which can be imaginary (CP violation) and large enough.

LHC limits are weakened,  $M_T \gtrsim 700 \,\mathrm{GeV}$ , due to dominant  $T \to St$  deca

Strongest constraints are from resonant *s* production + EWPD...

#### **Constraint from** *s* **production**

Strongest constraint is from  $gg \to s \to \gamma\gamma$ 





Constrains largest allowed value of  $\eta$ , hence of

$$\frac{1}{\Lambda} = \frac{y'\eta}{y_t M_T}$$

EWPD oblique parameter Tconstrains y' through  $t_L$ - $T_L$  mixing

# **EWPT & observable gravity waves**

A strongly first order transition can produce gravity waves, potentially observable by LISA experiment.

Mixing of s and h could modify hZZ coupling in a correlated way.



Orange: 1st order; blue: strongly 1st order (EWBG); green: very strongly 1st order (gravity waves)

But you need more than a strong EWPT for EWBG; you need to make large enough baryon asymmetry too!

# **Computing baryon asymmetry (BAU)**

Complex top-quark mass  $|m(z)|e^{i\theta(z)}$  in bubble wall produces a CP-violating force acting oppositely on particles/antiparticles:

$$\vec{F} = F\hat{z} \cong -\frac{(m^2)'}{2E} \pm \frac{(m^2\theta')'}{2E^2} + O(\theta^2) \quad [' = d/dz]$$

Distorts the phase space distribution of top quarks (and h,  $b_L$  that couple to t through collision term C) via Boltzmann equation

$$\left(\frac{d}{dt} + \vec{v} \cdot \vec{\nabla} + \frac{\vec{F}}{m} \cdot \vec{\nabla}_p\right) f = \mathcal{C}[f]$$

Split f into two pieces (JC, Joyce, Kainulainen hep-ph/0006119)

$$f_i(p, x) = \frac{1}{e^{\beta[\gamma_w(E + v_w p_z) - \mu_i(x)]} \pm 1} + \delta f_i(p, x)$$
  
deviation from chemical equilibrium  
encoded in  $\mu_i$ 

with form of  $\delta f_i$  unspecified. Linearize in perturbations  $\mu_i$ ,  $\delta f_i$ . Define "velocity potential"  $u_i = \int d^3 p (p_z/E) \delta f_i$ .

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# Fluid equations $\rightarrow$ BAU

Formalism developed by JC, Joyce, Kainulainen hep-ph/0006119, refined by Fromme, Huber hep-ph/0604159

In wall rest frame, linearized Boltzmann eq. (neglecting m' contributions) is



Take first two moments to derive fluid equations,

$$\int d^3 p \,(\text{B.E.}), \quad \int d^3 p \, \frac{p_z}{E} (\text{B.E.})$$

Demand that  $\int d^3 p \, \delta f_i \equiv 0$  ( $u_i$  is "orthogonal" to  $\mu_i$ )

Solve coupled eqs. for  $\mu_{t_L}$ ,  $\mu_{t_R}$ ,  $\mu_h$ ,  $\mu_{b_L}$ ; these feed into sphaleron rate equation,

$$\dot{n}_B = -\frac{3}{2} \frac{\Gamma_{\rm sph}}{T} \sum_{i=1}^3 \left( 3\mu_{u_L}^i + 3\mu_{d_L}^i \right)$$

Baryon asymmetry is integral of this.

#### The devil is in the details!

#### We do not make an explicit ansatz for the form of $\delta f$ .

From JC, Joyce, Kainulainen, hep-ph/0006119:

The  $\delta f(x, p)$  term will play the role of an auxiliary field in the following derivation. It cannot be neglected, for the Boltzmann equations with a force term can only be self-consistently solved if there is a perturbation which is anisotropic in momentum space.<sup>4</sup> The shape of  $\delta f_i(p, x)$  cannot be consistently restricted beyond (79), since this would require some hierarchy between the elastic interaction rates that would allow certain perturbations to be damped more slowly than others.

But we have to evaluate integrals like

$$\int d^3p \left(\frac{p_z}{E}\right)^2 \delta f \stackrel{?}{=} u \int d^3p \left(\frac{p_z}{E}\right) f_{0,v_w} \bigg/ \int d^3p f_{0,v_w}$$

with  $f_{0,v} = (e^{\gamma_w \beta (E+v_w p_z)} + 1)^{-1}$  in wall rest frame. We assume it *factorizes* as above.

This is necessary for getting the expected  $v_w$ -dependence. Adopting an explicit ansatz always leads to the wrong behavior.

# **Critical wall speed**

Define  $w = (\mu, u)$ ; fluid equations take matrix form

$$\underbrace{Aw' - (m^2)'Bw}_{} = \underbrace{S}_{} + \underbrace{Cw}_{}$$

Liouville terms



with

$$A = \begin{pmatrix} v_w K_1 & 1 \\ -K_4 & -v_w \end{pmatrix}, \qquad K_i = K_i(v_w, m/T)$$

Diffusion breaks down if det  $A \rightarrow 0$ ; this defines critical wall speed  $v_c$ :



Common approximation: evaluate  $K_i$  at  $v_w = 0$ . At m/T = 0,  $K_1 \rightarrow 1, K_4 \rightarrow 1/3, v_c = 1/\sqrt{3}$ .

This may have reinforced the "lore"

But if we keep the full  $v_w$ -dependence of  $K_i$ , we get the right answer,  $v_c = 1!$ 

## **Explicit ansatz problem**

The first papers did choose explicit form for  $\delta f$  (Joyce, Prokopec, Turok hep-ph/9410282; Moore, Prokopec hep-ph/9506475):

$$f(\vec{p}, \vec{x}, t) = rac{1}{e^{eta(\gamma(E - ec{v} \cdot ec{p})) - \mu} \pm 1}$$

with three perturbations:  $w = (\mu, \delta T/T, v)$ , In the  $m/T \rightarrow 0$  limit, the A matrix is

$$A \equiv \begin{pmatrix} v_w c_2 & v_w c_3 & \frac{1}{3}c_3 \\ v_w c_3 & v_w c_4 & \frac{1}{3}c_4 \\ \frac{1}{3}c_3 & \frac{1}{3}c_4 & \frac{1}{3}v_w c_4 \end{pmatrix}; \ c_2 = \frac{1}{6}, \ c_3 = \frac{9\zeta(3)}{4\pi^2}, \ c_4 = \frac{7\pi^2}{60}$$

Singular at  $v_w = 1/\sqrt{3}$ ! Origin of the "lore."

This is still the standard practice for computing CP-even perturbations, needed to find properties of the bubble wall (friction, wall speed and shape).

#### **Full fluid equation network**

# We correct the results of Fromme, Huber hep-ph/0604159; most coefficient functions $K_i$ are revised, some new ones added



#### The baryon asymmetry

Integrating the sphaleron rate equation gives

$$\eta_B = \frac{405\,\Gamma_{\rm sph}}{4\pi^2 v_w \gamma_w g_* T} \int dz\,\mu_{B_{\rm L}} f_{\rm sph}\,e^{-45\Gamma_{\rm sph}|z|/4v_w \gamma_w}$$

with strong sphalerons relating  $\mu_{u,d,s,c}$  to  $\mu_{t,b}$ :

 $\mu_{B_L} = \frac{1}{2} (1 + 4D_0^t) \mu_{t_L} + \frac{1}{2} (1 + 4D_0^b) \mu_{b_L} + 2D_0^t \mu_{t_R}.$ 



# **Other comparisons**



#### Conclusions

EWBG is a testable framework for baryogenesis. LHC continues to put pressure on allowed models, but possiblities still exist.

Strong 1st order transitions naturally favor high wall speeds, which also tend to produce observable gravity waves.

Improved transport equations make EWBG feasible for supersonic walls, although challenging to get large enough BAU

These improvements should be applied to the CP-even fluid perturbations, to see how they may effect the microscopically computed wall properties (especially  $v_w$  and  $L_w$ )

Collision terms should be reevaluated. Some date back to 1996!

#### **Extra slides**

# LHC limit on top partner

ATLAS and CMS constraint vectorlike top-quark partner mass > 1 TeV if  $T \rightarrow Zt$  or  $T \rightarrow ht$  dominates. In our model,  $T \rightarrow st$  dominates, and  $s \rightarrow gg$  is main singlet decay channel.

