



Polarimetry for Axion Search

Qazal Rokn

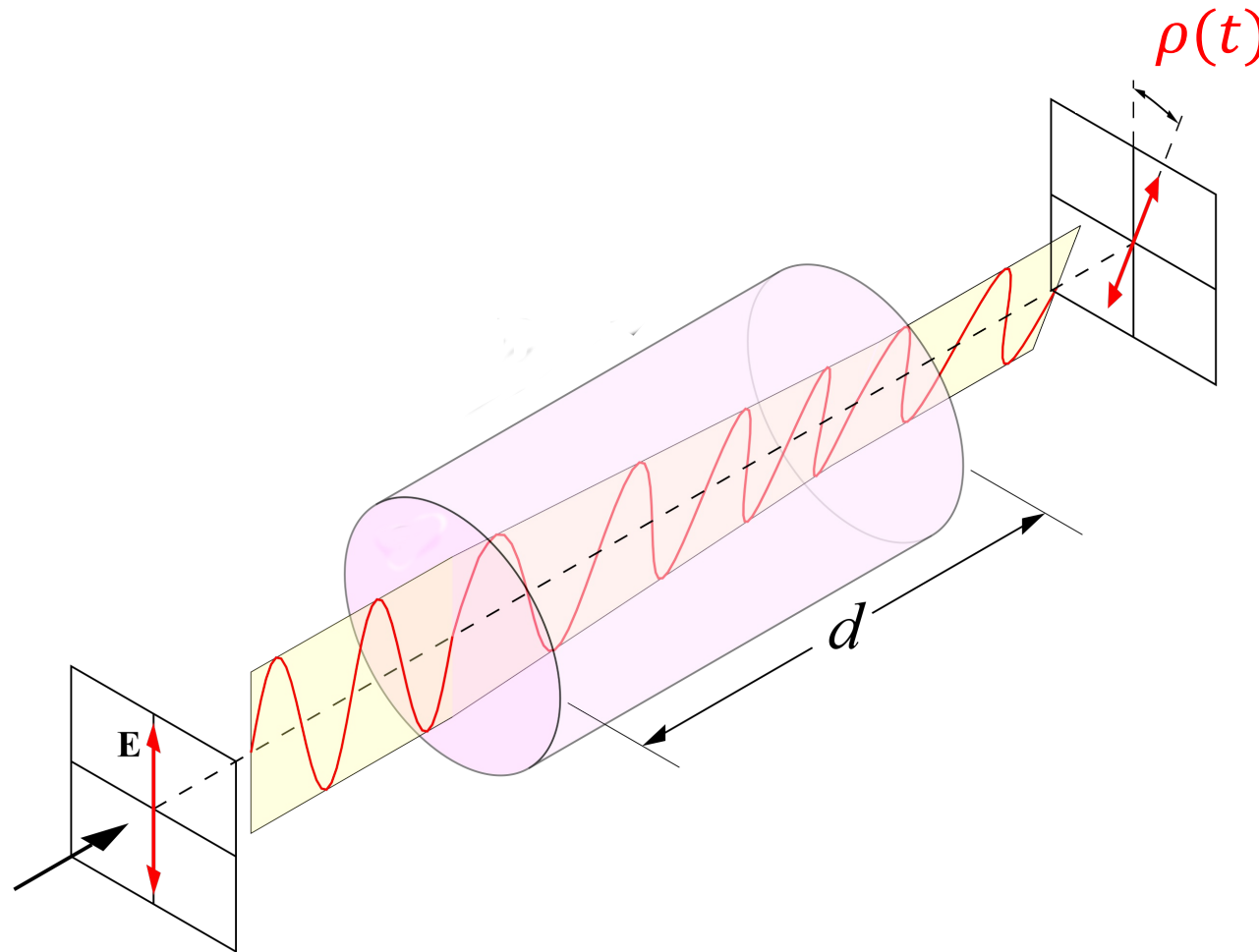
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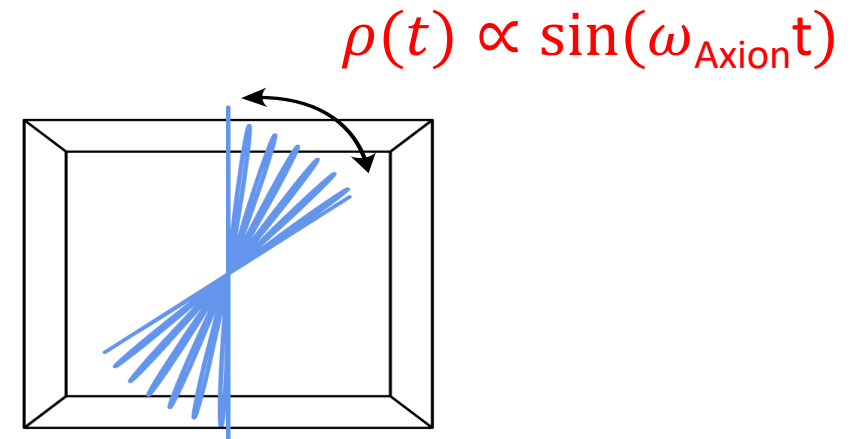
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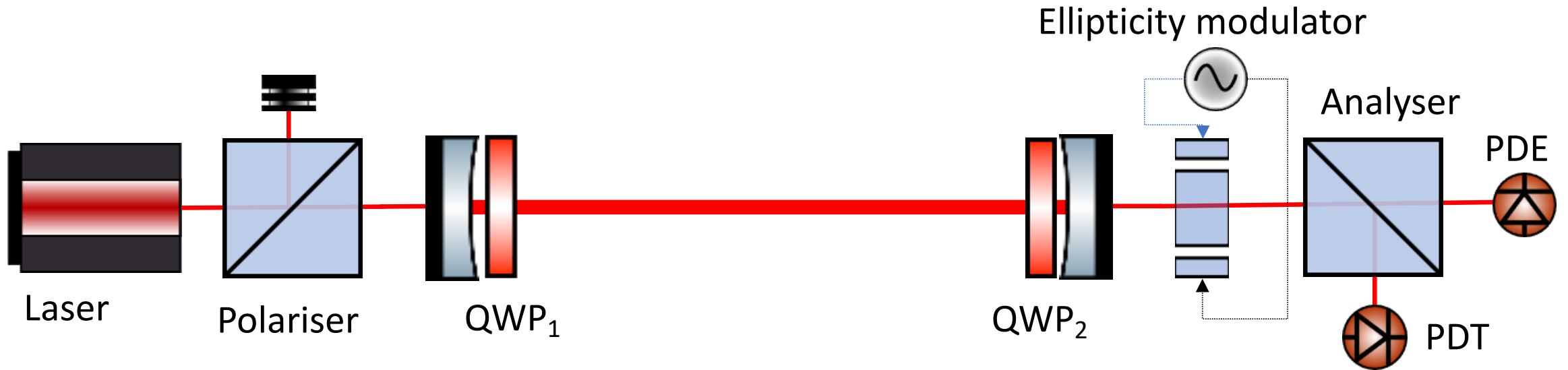
Effect of Axion Field on Polarization of Light



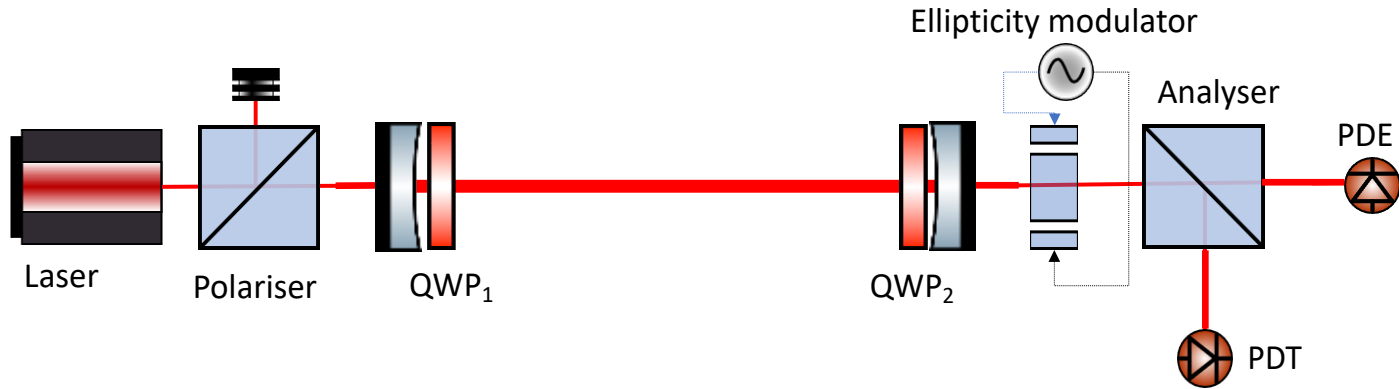
- Axion field rotates the polarization of linearly polarized light
- Angle of rotation oscillates with the frequency of Axion field



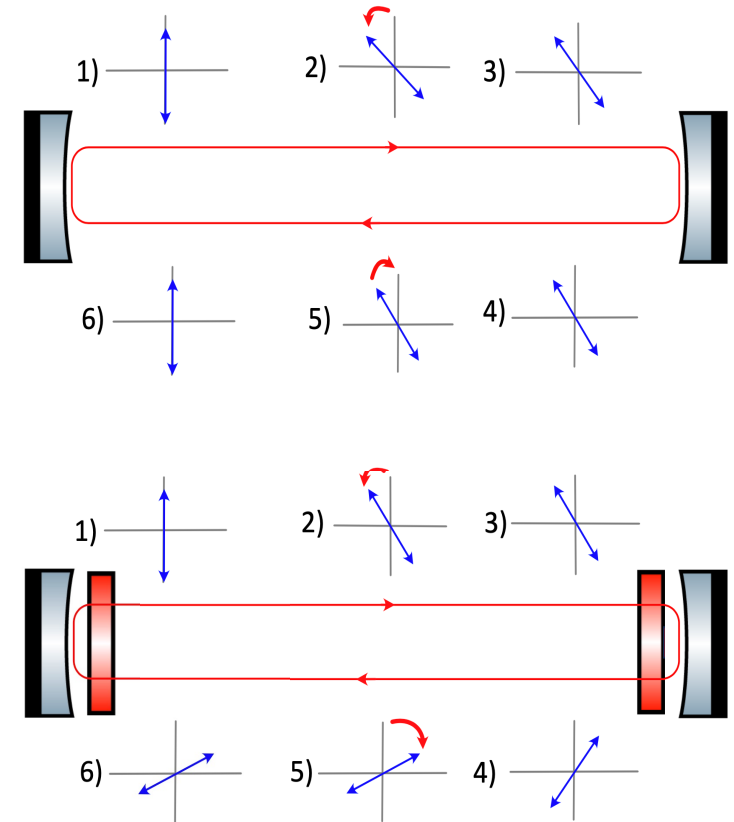
Proposed set up for Axion Detection



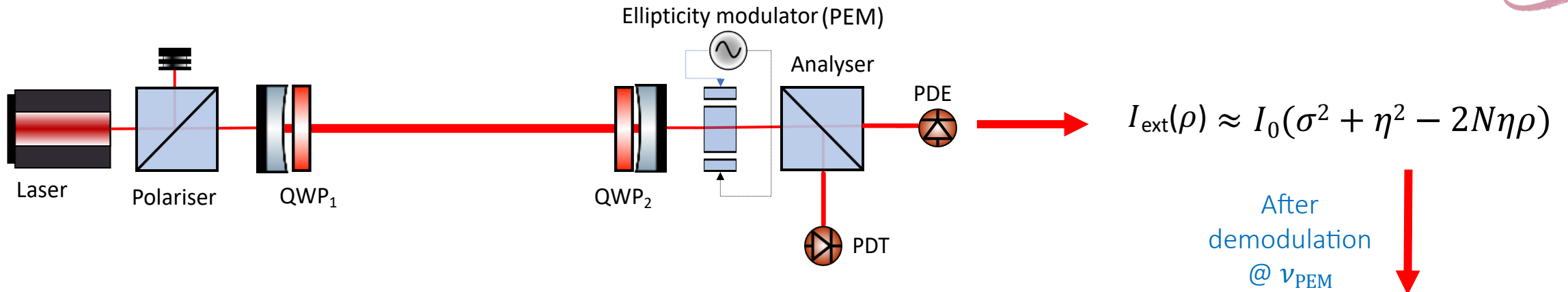
Function of the Quarter Wave Plates



- Axion-induced rotation direction depends on propagation direction
- The quarter wave plates prevent cancellation of the Axion effect
- QWP converts rotation modulation to ellipticity modulation



Detection using Heterodyne Technique

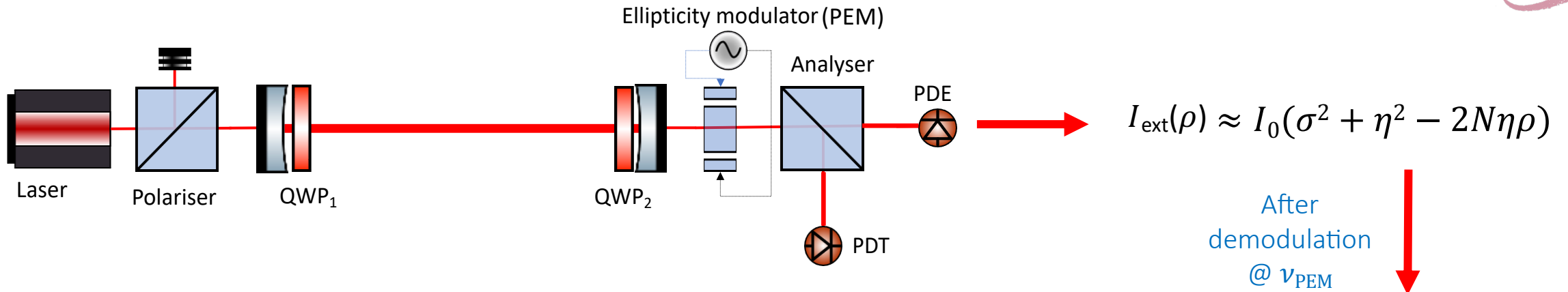


- Ellipticity-modulated light exits the cavity
- The light then passes through a Ellipticity Modulator ($\eta = \eta_0 \cos(\omega_{\text{PEM}} t)$)
- The resulting beat note is linearly proportional to the axion-induced rotation angle

$$|\rho(\nu)| = \frac{\mathbb{I} \nu_{\text{PEM}}}{N h_T I_0 \eta_0}$$

Symbol	Definition
ρ	Rotation angle caused by the axion
η	PEM modulation depth
σ	Extinction ratio
$N = 2\mathcal{F}/\pi$	Cavity buildup
I_0	The intensity of light leaving the cavity

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$$\frac{\text{Noise} (I_{\nu_{PEM}})}{I_0 \eta_0} :$$

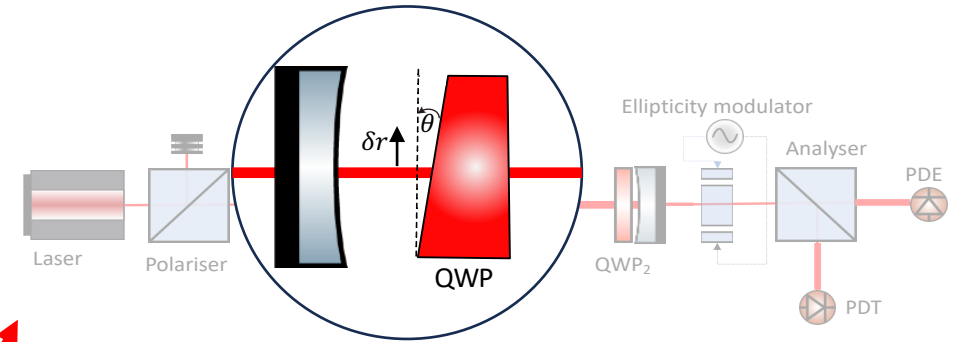
- Shot noise = $\sqrt{\frac{2e}{qI_0} \left(\frac{\sigma^2 + \eta_0^2/2}{\eta_0^2} \right)}$
- Dark noise = $\frac{i_{\text{dark}}}{qI_0 \eta_0}$
- Thermal Johnson noise = $\sqrt{\frac{4k_B T}{G}} \frac{1}{qI_0 \eta_0}$
- RIN = $N_{\nu_m}^{(\text{RIN})} \frac{\sqrt{(\sigma^2 + \eta_0^2/2)^2 + (\eta_0^2/2)^2}}{\eta_0}$
- Seismic/ pointing noise = $2\pi \frac{N \delta r \theta}{\lambda} \Delta n \gamma(L, f) / I_0 \eta_0$

Noise Budget



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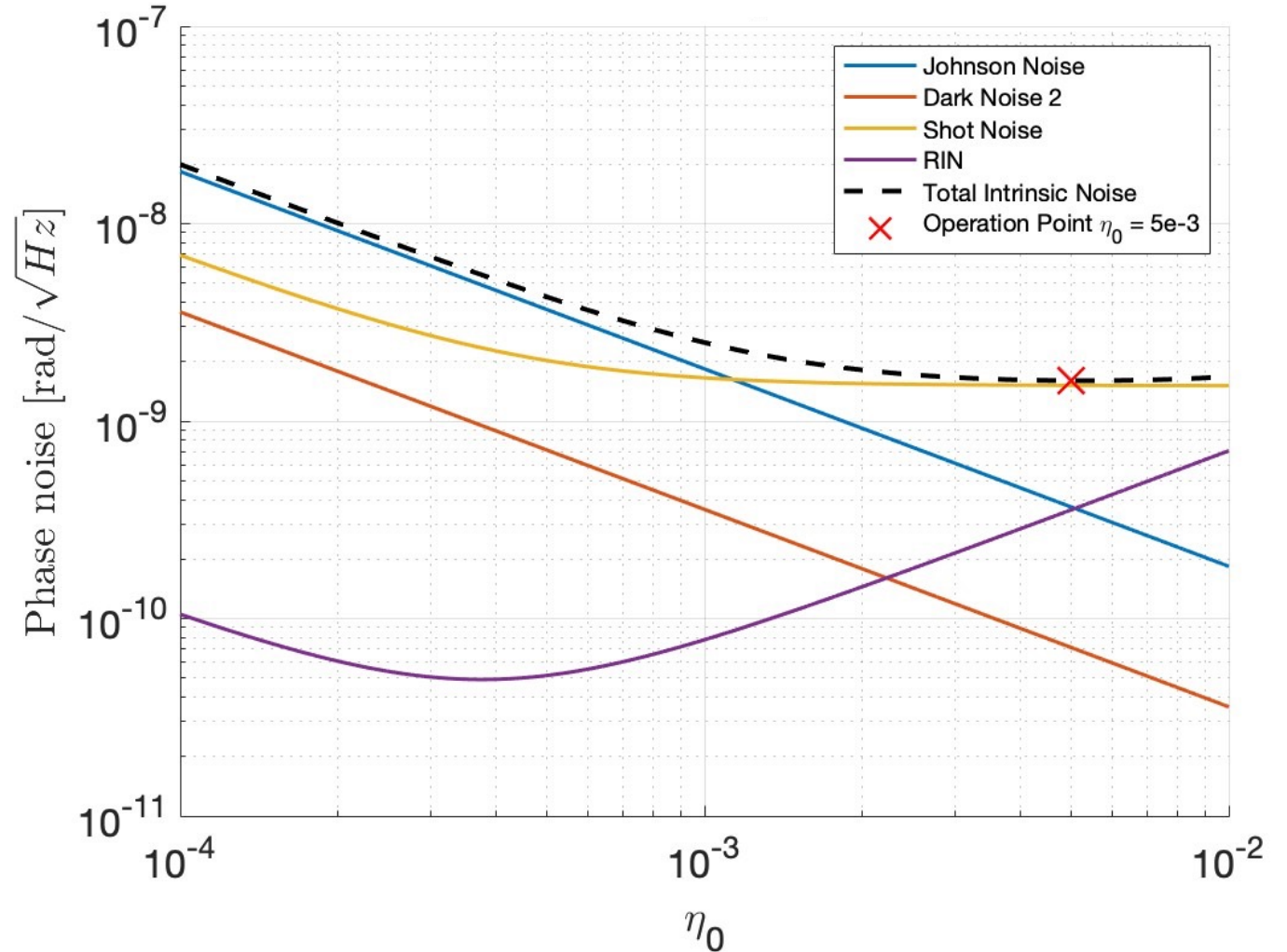


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- RIN = $N_{\nu_m}^{(RIN)} \frac{\sqrt{(\sigma^2 + \eta_0^2/2)^2 + (\eta_0^2/2)^2}}{\eta_0}$

Symbol	Value
I_0	100 mW
σ^2	1e-7
T	300 k
G	1e6 Ω
q	0.7 A/W
$N_{\nu_m}^{(RIN)}$	1e-7/ $\sqrt{\text{Hz}}$
i_{Dark}	25 fA _{rms} / $\sqrt{\text{Hz}}$



Sensitivity to the Axion-photon Coupling Coefficient



T	Transmittivity
R	Reflectivity
τ	Cavity round trip time
ρ_{local}	Dark matter local density
N	Cavity buildup
\mathcal{F}	Finesse

$$|\rho(\nu)| = \frac{S_P^{(tot)}}{N h_T(\nu)}$$

cavity build up: $N = \frac{2\mathcal{F}}{\pi} \approx \frac{2}{T + \text{loss}}$

$h_T(\nu) = \frac{T}{\sqrt{1 + R^2 - 2R \cos 2\pi\nu\tau}}$: Cavity transfer function

Sensitivity to ρ : $|\rho(\nu)| \approx S_p^{tot} \frac{\sqrt{\text{loss}^2 + 4\sin^2(\pi\nu_{Axion}\tau)}}{2}$

$\rho \approx g_{a\gamma} a_0 \omega_a \tau \sin(\omega_a t)$

Sensitivity to $g_{a\gamma}$: $g_{a\gamma} = \frac{S_p^{tot}}{2\tau} \sqrt{\frac{\text{loss}^2 + 4\sin^2(\pi\nu_{Axion}\tau)}{2\rho_{local}}}$

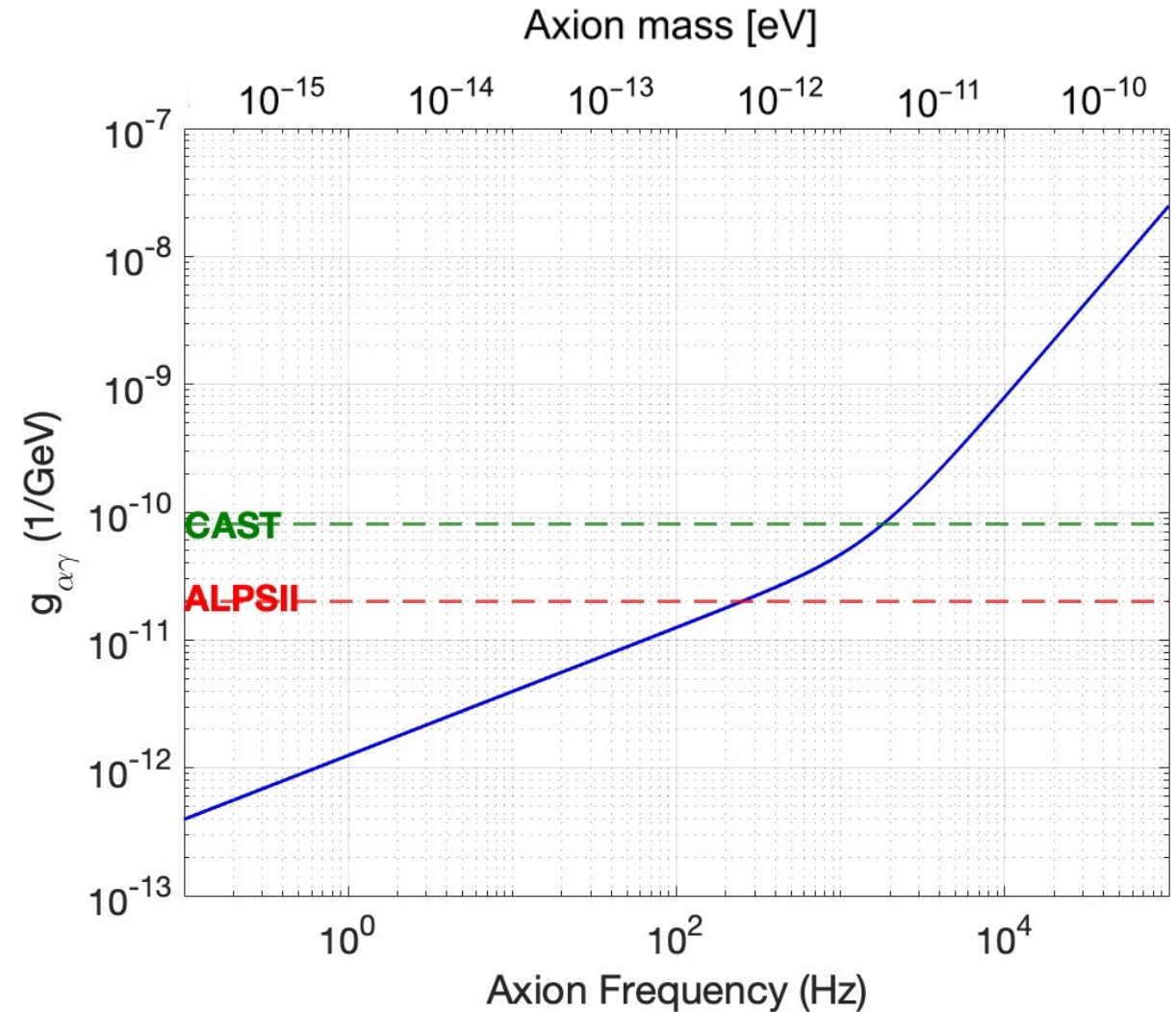
Sensitivity to the Axion-photon Coupling Coefficient



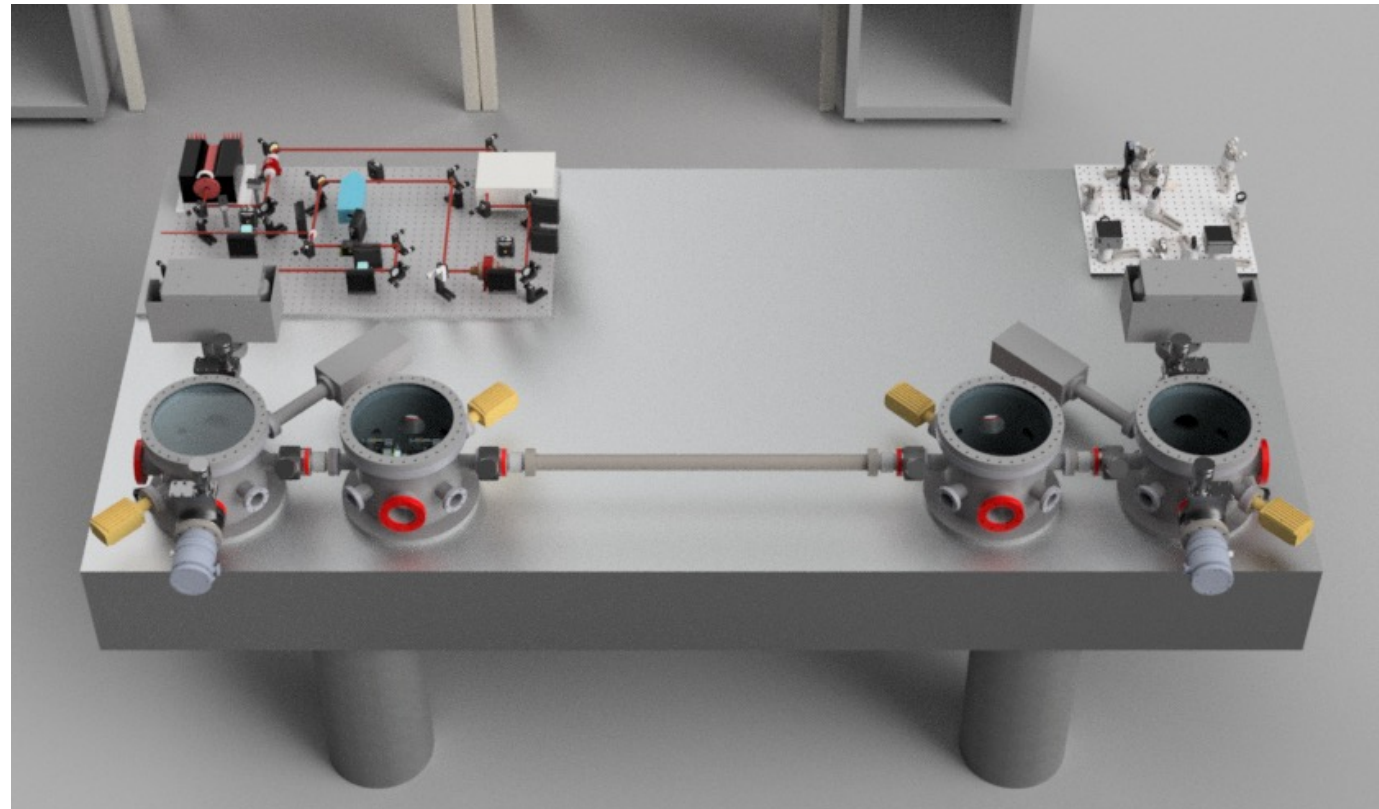
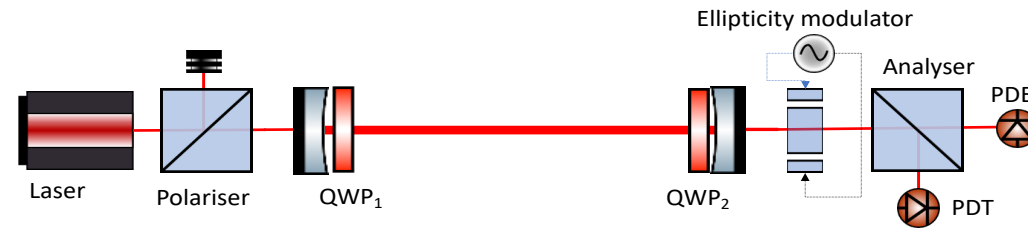
- The current sensitivity analysis does not include the effect of seismic noise

$$g_{\alpha\gamma} = \frac{S_p^{tot}}{2\tau} \sqrt{\frac{Loss^2 + 4\sin^2(\pi\nu_{Axion}\tau)}{2\rho_{local}}}$$

Symbol	Value
I_0	100 mW
σ^2	1e-7
T	300 k
G	1e6 Ω
q	0.7 A/W
$N_{v_m}^{(RIN)}$	1e-7/ $\sqrt{\text{Hz}}$
i_{Dark}	25 fA _{rms} / $\sqrt{\text{Hz}}$
Loss	100 ppm
N	20,000



3D Design of Experimental Setup





Thank you!

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