



Axion electrodynamics in plasmas

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Presentation Overview

- 1 Overview of axion-electrodynamics
- 2 Plasmas and axions
- 3 Polariton quantization

Theoretical basis

Genesis of Axion theory

- 1 Axial U(1) problem
- 2 QCD vacuum angle
- 3 Strong CP problem
- 4 PQ symmetry

The axion is represented by a spin-zero (boson) scalar field φ ,

$$\mathcal{L}_{\text{eff}} = \underbrace{\frac{1}{2}\partial_\mu\varphi\partial^\mu\varphi - \frac{1}{2}m_0^2\varphi^2}_{\text{free axion}} + \underbrace{\mathcal{L}_{\varphi\gamma}}_{\text{axion-photon}} + \underbrace{\mathcal{L}_{\varphi\psi}}_{\text{axion-fermion}} + \dots \quad (1)$$

Axion-photon coupling

The axion couples to the electromagnetic field via:

$$\mathcal{L}_{\varphi\gamma} = \frac{\kappa}{4\mu_0} \varphi F_{\mu\nu} \tilde{F}^{\mu\nu} \quad (2)$$

$\kappa = \eta m_0$ is the coupling constant. η [GeV⁻²] model dependent.

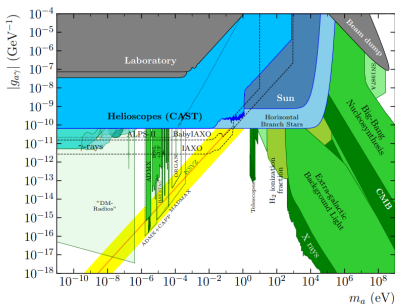


Figure: [Irastorza 2022] Sensitivity of the axion-detection experiments and cosmological bounds (green) in the coupling-mass parameter space. The QCD axion is shown in yellow.

Axion electromagnetism

The axion-photon coupling $\mathcal{L}_{\varphi\gamma}$ effectively rotates the fields (\mathbf{E} , \mathbf{B}) by an angle $\arctan(-\kappa\varphi)$ in the Maxwell equations [Visinelli 13]

$$\nabla \times (\mathbf{E} - c\kappa\varphi\mathbf{B}) = -\frac{\partial}{\partial t}(c\mathbf{B} + \kappa\varphi\mathbf{E}) \quad \nabla \cdot (\mathbf{E} - \kappa\varphi\mathbf{B}) = \frac{\rho}{\varepsilon_0} \quad (3)$$

$$\nabla \times \left(\mathbf{B} + \frac{\kappa\varphi}{c}\mathbf{E} \right) = -\frac{1}{c^2} \frac{\partial}{\partial t} (\mathbf{E} - c\kappa\varphi\mathbf{B}) + \mu_0\mathbf{J} \quad \nabla \cdot \left(\mathbf{B} - \frac{\kappa\varphi}{c}\mathbf{E} \right) = 0 \quad (4)$$

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The propagation for the axion field is given by a KG eq^o

$$\left(\square + m_\varphi^2 c^2\right)\varphi = \frac{2g}{\mu_0} \underbrace{\mathbf{E} \cdot \mathbf{B}}_{\propto F_{\mu\nu}\tilde{F}^{\mu\nu}} \quad (5)$$

Photon wave equation

Axion-Maxwell equations with dual symmetry:

$$\partial_\mu (R_\varphi \mathcal{F}^{\mu\nu}) = \mu_0 \mathcal{J}^\nu$$

$\mathcal{F}^{\mu\nu} = (F^{\mu\nu}, F^{*\mu\nu})^T$ dual vector; $F^{\mu\nu} \equiv \partial^\mu A^\nu - \partial^\nu A^\mu$;

$R_\varphi = \begin{pmatrix} 1 & -\kappa\varphi \\ \kappa\varphi & 1 \end{pmatrix}$ axion rotation matrix in dual space.

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Axion-photon dual wave equation:

$$\partial^\mu \partial_\mu \tilde{A}^\nu - \partial^\nu \partial_\mu \tilde{A}^\mu = \mu_0 \mathcal{J}^\nu$$

$\tilde{A}^i = A^i + X^i$, with gauge offset $\partial^i \times X^i = \kappa\varphi E^i$, $\partial_0 X^i + \partial^i X_0 = \kappa\varphi B^i$

Axion-Polariton gauge

Lorenz gauge for axion-polariton in vacuum

$$\partial_\mu \tilde{\mathcal{A}}^\mu = 0$$

Photon mass and axion-photon mass can not be simultaneously eliminated by a change of gauge.

This choice ensures usual current conservation $\partial_\mu j^\mu = 0$ and vanishing longitudinal axion-polariton via Ward-Takahashi

$$k_\mu \tilde{\epsilon}^\mu = 0$$

Note: fixing instead $\partial_\mu A^\mu = 0$, conserved current is $j^\nu - j_{ax}^\nu$,

$$j_{ax}^\nu = -\partial_\mu (\kappa \varphi) F^{*\mu, \nu}$$

Photon angular momentum

Effect on photon angular momentum:

$$\mathbf{S} = \frac{1}{c} \int d^3x \pi_{\mathbf{A}} \times \mathbf{A} \quad (6)$$

$$\pi^\alpha = \frac{\partial L}{\partial(\partial_0 \mathbf{A}_\alpha)} = -\mathbf{E} + \kappa\varphi \mathbf{B}. \rightarrow \mathbf{S} = \mathbf{S}_0 + \mathbf{S}_\chi$$

$$\mathbf{S}_\chi = \frac{1}{c} \int d^3x \kappa\varphi (\nabla \times \mathbf{A}) \times \mathbf{A} = 0 \quad (7)$$

Axion coupling contribution to photon spin is vanishing \Rightarrow no axion-polariton in vacuum (states are degenerate)

$(\mathbf{J}_Z^{ax} = 0 \leftrightarrow \mathbf{J}_Z^\gamma = \pm 1)$ impossible from longitudinal field alone.

Some more context

[Raffelt & Stodolsky]: mixing possible with QED vacuum polarization.

also [Raffelt & Stodolsky]: don't forget to add *neutral gas* contribution to birefringence effect for realistic astrophysical applications.

Plasma (the realistic astrophysical application):



*To take with a grain of salt

Plasmas → Langmuir Waves (plasmons)

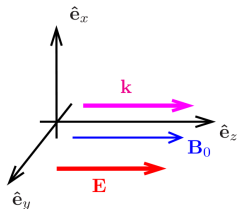


Figure: In plasmas, resistive effects typically induce $\mathbf{E} \cdot \mathbf{B} \neq 0$. A small displacement $\delta z = \xi$ of an electron from its equilibrium position induces oscillations:

$$\frac{d^2\xi}{dt^2} + \omega_{pe}^2\xi = 0, \quad \omega_{pe} = \frac{4\pi n_e e^2}{m_e}$$

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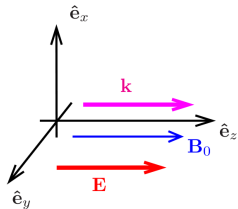


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At finite temperatures, $k \neq 0$,

$$\omega_{pl}^2 = \omega_{pe}^2 + S_e^2 k^2, \text{ plasmon bare disp.} \quad (8)$$

$$\omega_{\varphi}^2 = m_{\varphi}^2 c^4 + c^2 k^2, \text{ axion bare disp (KG)} \quad (9)$$

(8): $S_e = \sqrt{\frac{3k_b T_e}{m_e}}$ thermal speed

(9): $m_{\varphi} = \sqrt{m_0^2 + 2g^2 B_0^2 / (c\mu_0)}$ axion **effective** mass

Axion-plasmon Waves

Equations (4) and (5) must be closed with continuity equations within a fluid treatment:

$$\frac{\partial n_e}{\partial t} + \nabla \cdot (n_e \mathbf{u}) = 0 \quad (10)$$

$$\left(\frac{\partial}{\partial t} + \mathbf{u} \cdot \nabla \right) \mathbf{u} = \frac{\nabla P}{m_e n_e} - \frac{n_e}{m_e} (\mathbf{E} + \mathbf{u} \times \mathbf{B}). \quad (11)$$

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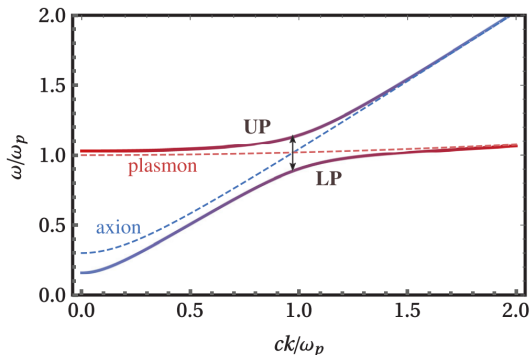
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Yields a coupled dispersion relation on $(\tilde{n}, \tilde{\varphi})$:

$$(\omega^2 - \omega_{pl}^2) \tilde{n} - i \frac{g e c}{m_e} B_0 k \tilde{\varphi} \quad (12)$$

$$(-\omega^2 + \omega_{\varphi}^2) \tilde{\varphi} + i \frac{\kappa e c^2}{\epsilon_0 k} B_0 \tilde{n} = 0 \quad (13)$$

Axion-plasmon polaritons [Terças+18]



$$\omega_{U,L}^2 = \frac{1}{2} \left(\omega_\varphi^2 + \omega_{pl}^2 \pm \sqrt{(\omega_{pl}^2 - \omega_\varphi^2)^2 + 4\Omega^4} \right) \quad (14)$$

$$\Omega = (2\kappa^2 e^2 B_0^2 (n_0/m_e) c^5 / \hbar)^{1/4} \text{ Rabi frequency}$$

Rotating wave approximation

The system can be simplified (approximated) near the resonance $\omega_{pl} \simeq \omega_p$, with $\omega \rightarrow -i \frac{\partial}{\partial t}$:

$$\begin{aligned} \left(i \frac{\partial}{\partial t} - \omega_{pl} \right) \tilde{n} - \frac{igec^4 B_0}{\hbar k \omega_p} \tilde{\varphi} &= 0 \\ \left(i \frac{\partial}{\partial t} - \omega_{\varphi} \right) \tilde{\varphi} + \frac{igec^4 B_0}{\hbar k \omega_p} \tilde{n} &= 0 \end{aligned}$$

Axion-plasmon quantization:

$$\tilde{n}(x) = \sum_k \mathcal{A}_k \left(\hat{a}_k e^{ikx} + h.c \right) \tilde{\varphi}(x) = \sum_k \mathcal{B}_k \left(\hat{b}_k e^{ikx} + h.c \right) \quad (15)$$

Axion-plasmon quantization

Heisenberg picture: $\frac{d\hat{c}}{dt} = \frac{i}{\hbar}[\hat{H}, \hat{c}]$,

$$\hat{H}_{\gamma\varphi} = \sum_k \omega_{pl} \hat{a}_k^\dagger \hat{a}_k + \sum_k \omega_\varphi \hat{b}_k^\dagger \hat{b}_k + \Omega \sum_k \hat{a}_k^\dagger \hat{b}_k + h.c. \quad (16)$$

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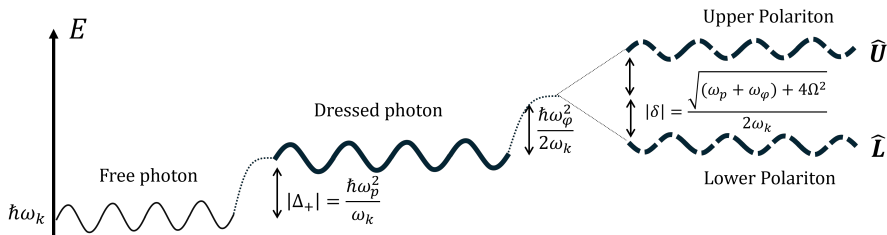
Mixing terms $\propto \Omega$ can be absorbed in diagonalization with polariton operators $\hat{L}_k = u_k \hat{a}_k - v_k \hat{b}_k$ and $\hat{U}_k = u_k \hat{b}_k + v_k \hat{a}_k$

$$\hat{H}'_{\gamma\varphi} = \tilde{\omega}_L \sum_k \hat{L}_k^\dagger \hat{L}_k + \tilde{\omega}_U \sum_k \hat{U}_k^\dagger \hat{U}_k \quad (17)$$

$$\tilde{\omega}_{U,L} = \frac{1}{2} \left(\omega_\varphi + \omega_{pl} \pm \sqrt{(\omega_{pl} - \omega_\varphi)^2 + 4\Omega^2} \right) \quad (18)$$

Conclusion

- $\cancel{U(1)}_{PQ} + \mathbf{B}_0 \xrightarrow[\text{mixing}]{\gamma_{\parallel} \leftrightarrow \varphi} \cancel{SO(2)}_{EM}$ axion-polariton level splitting



- Opens up much phenomenology: decay rates $\Gamma_{\varphi \rightarrow \gamma\gamma}$ [Terças+18], Boson oscillation [Capolupo 19], Casimir effect [Brevik 22], etc.

Discussion

- couplings neglected:
 - 1 $\mathcal{L}_{\psi\gamma}$: invitation to full quantization of AQED.
 - 2 \mathcal{L}_{EH} : can be added, but is it wise?