

Probing freeze-in via invisible Higgs decay

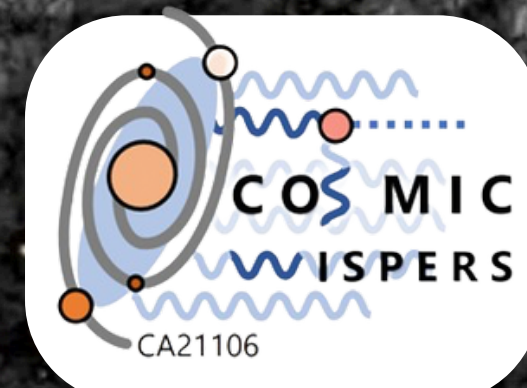
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Aveiro University - Portugal

2nd Training School of the COST Action: Cosmic WISPers (CA21106)

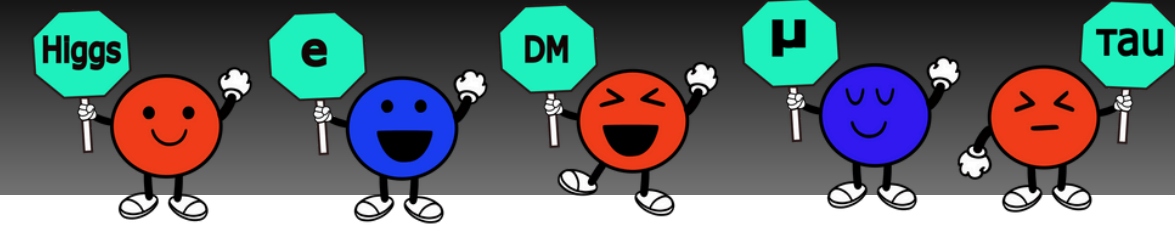


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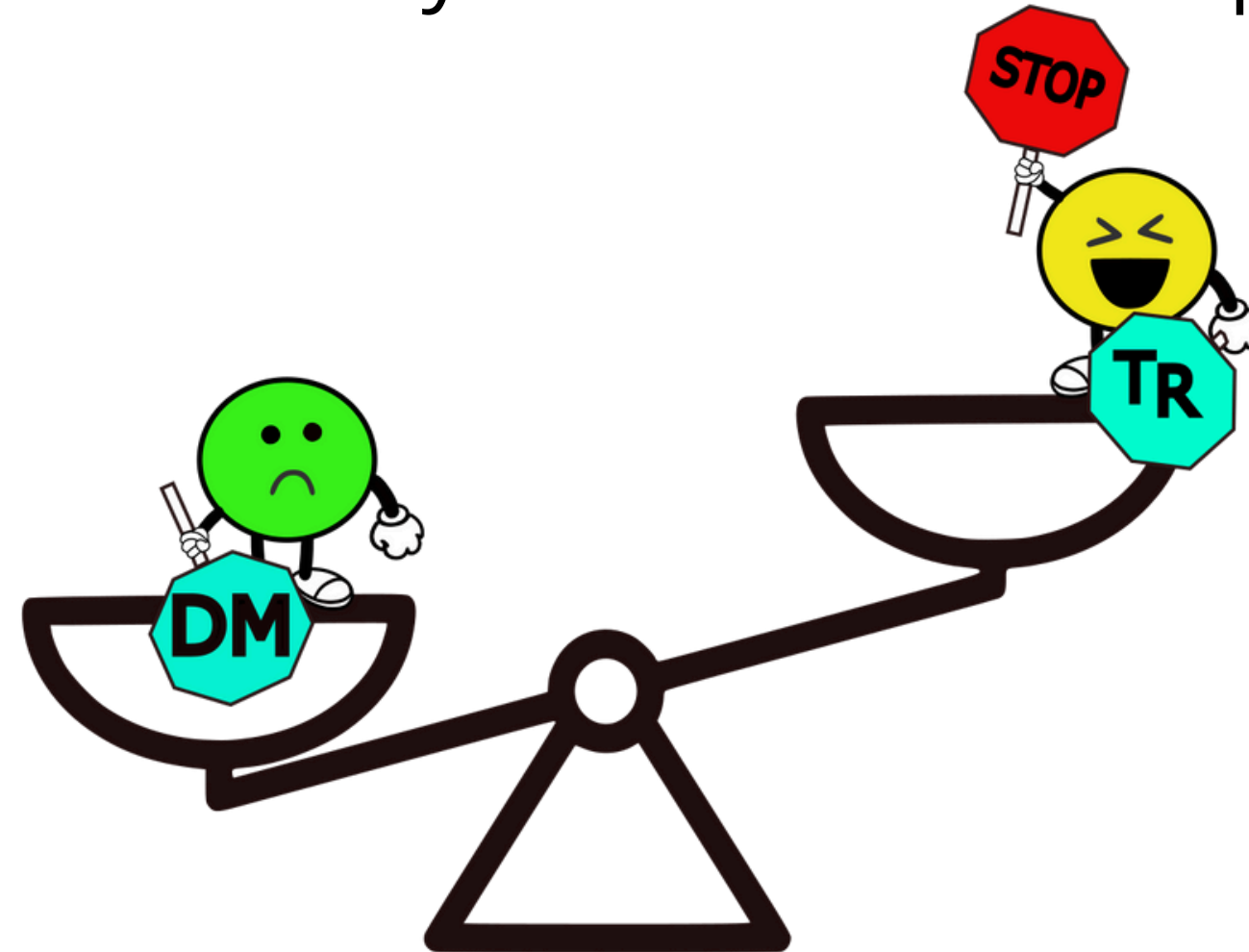
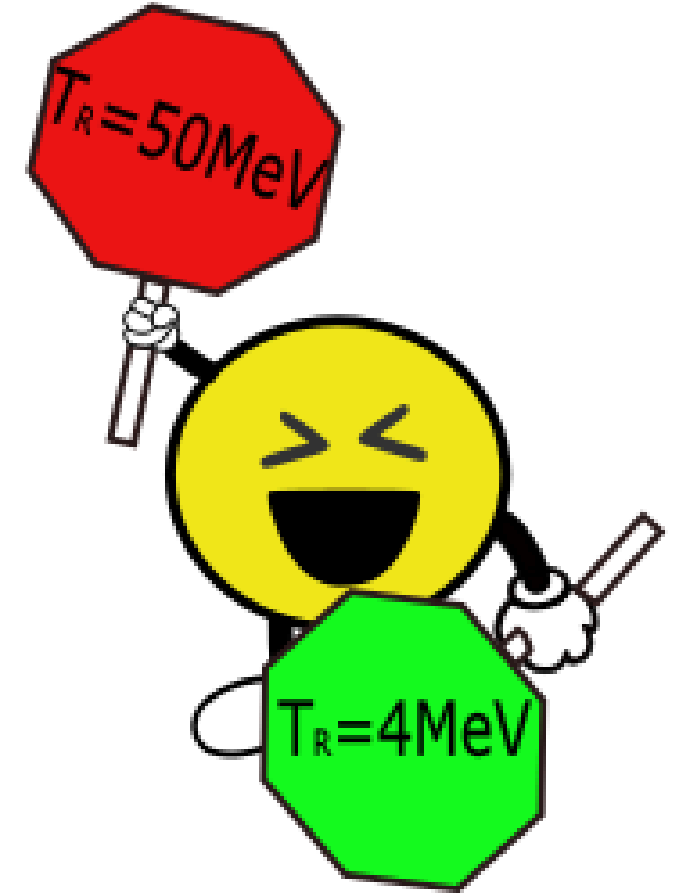


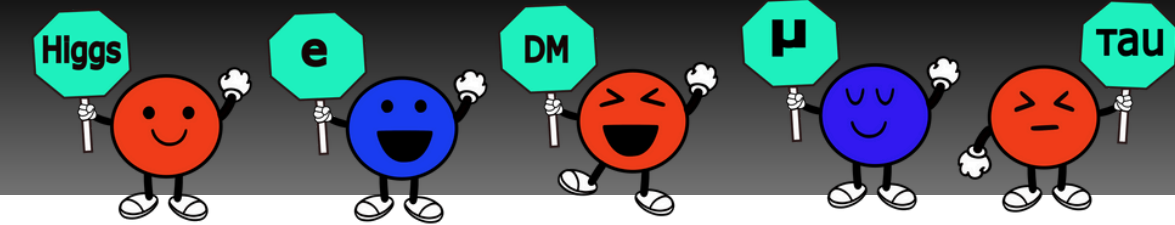
Fundação
para a Ciência
e a Tecnologia

Goal



- Study DM production in a low-reheating temperature $T_R \sim O(4 - 50)MeV$ scenario;
- We are focus on Higgs portal scenario;
- We are interested in the case where $m_{DM} \gg T_R$;
- The DM is produced non-thermally due Boltzmann suppression.



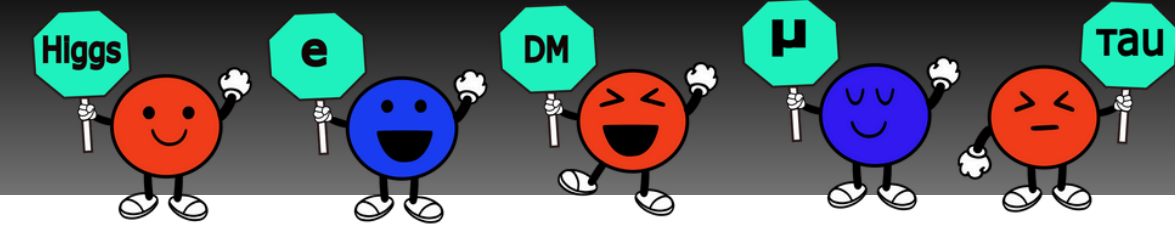


- We assume that the scalar S DM candidate couples with SM particles via the Higgs portal, with the Lagrangian

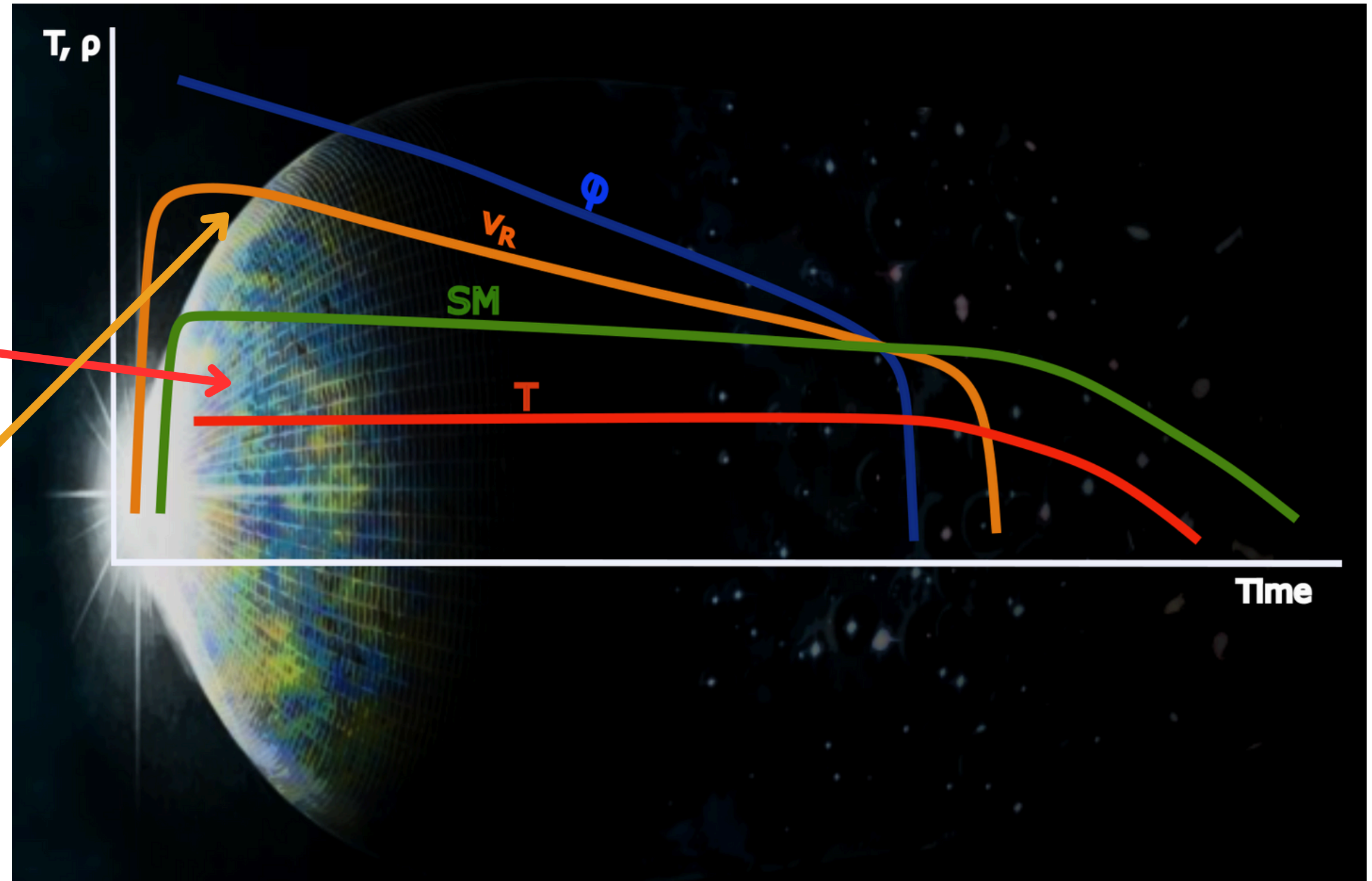
$$L \supset \frac{gm_f}{2m_W} h \bar{f} f + \frac{\lambda_{hs} v_{SM}}{2} h S S, \quad \rightarrow \text{Dark Side}$$

\rightarrow Standard Higgs interaction with fermions

where f represents a fermionic SM particle.



- We will assume that the maximum temperature of Universe is $T_{MAX} \sim T_R$.
- We can assume it for the scenario where the inflaton decays predominantly in other particle species (e.g. sterile neutrinos).



Boltzmann Equation



- Boltzmann equation for number density
Counts production / annihilation of 2 identical particles

$$\dot{n} + \underline{3Hn} = \underline{2\Gamma (ff \rightarrow SS)} - \underline{2\Gamma (SS \rightarrow ff)}$$

Production term Annihilation term

Dilution effect of Universe expansion

Boltzmann Equation



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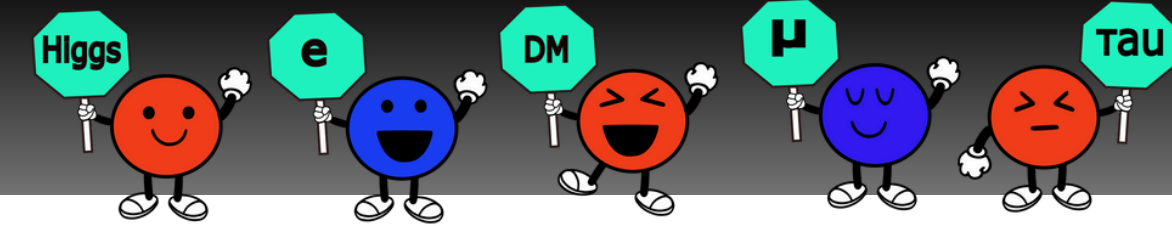
Production term
Annihilation term

Dilution effect of Universe expansion

$$\Gamma (XX \rightarrow YY) \equiv \langle \sigma_{XX \rightarrow YY} v \rangle n_X^2$$

Thermally averaged cross section

Boltzmann Equation



- Boltzmann equation for number density

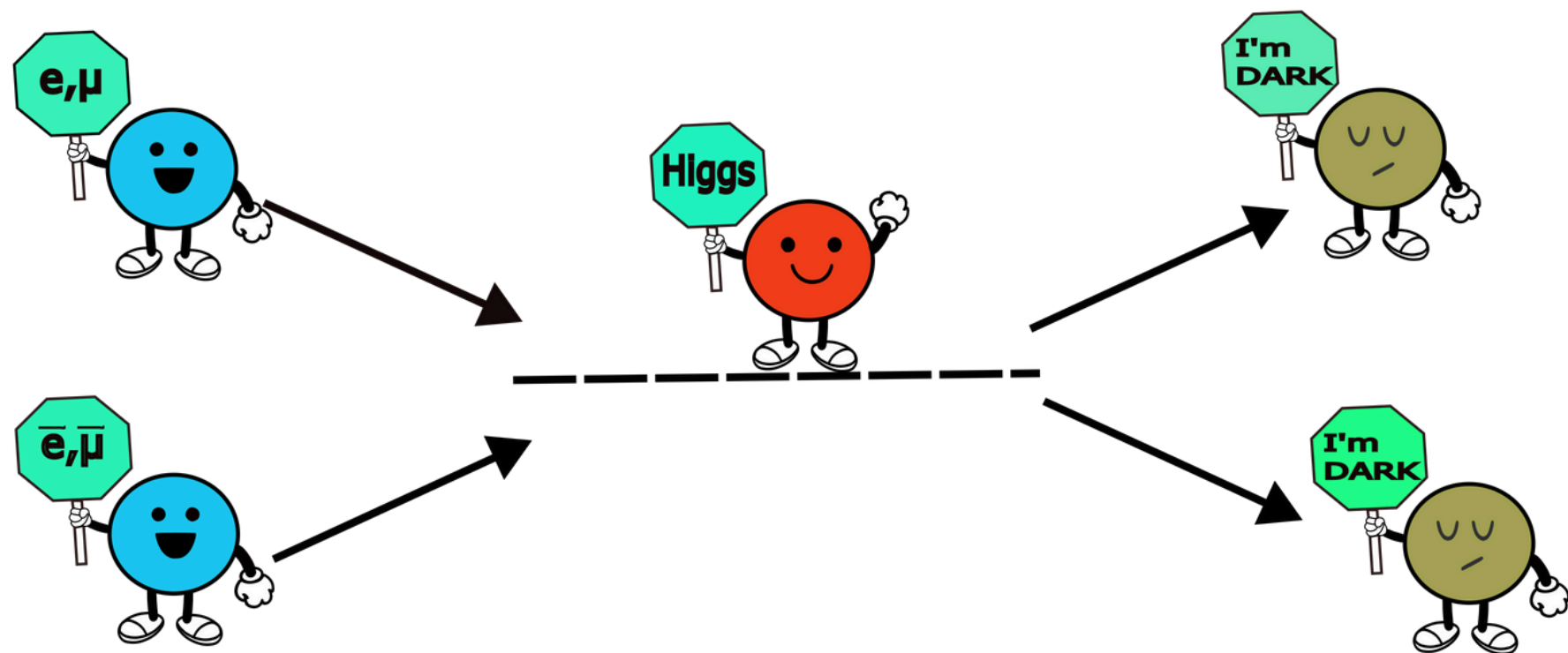
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$$\dot{n} + \underbrace{3Hn}_{\text{Dilution effect of Universe expansion}} = \underbrace{2\Gamma(ff \rightarrow SS)}_{\text{Production term}} - \underbrace{2\Gamma(SS \rightarrow ff)}_{\text{Annihilation term}}$$

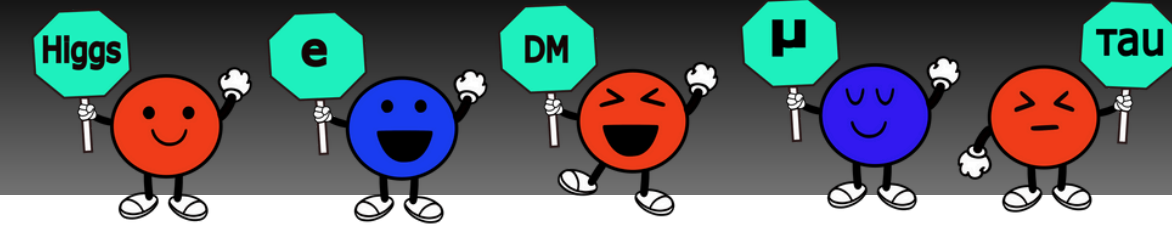
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Thermally averaged cross section

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Boltzmann Equation



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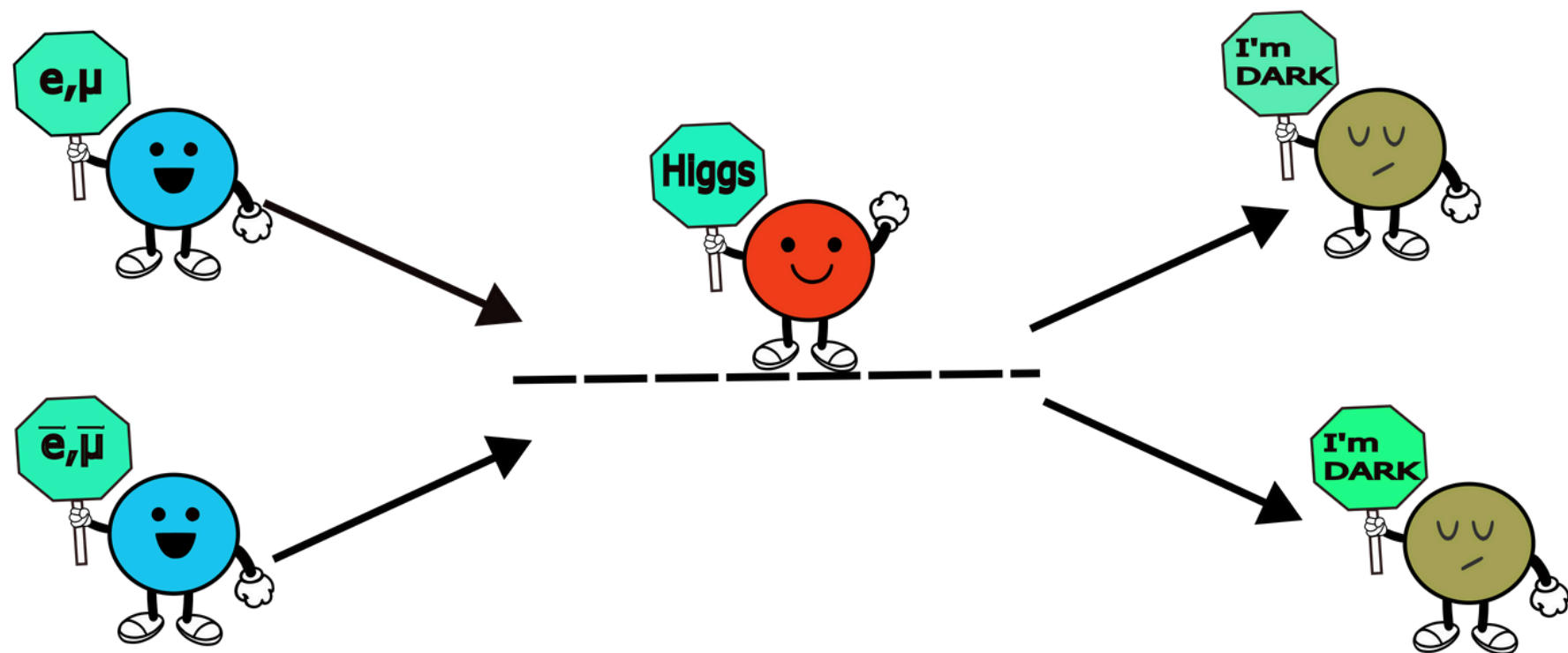
Production term
Annihilation term

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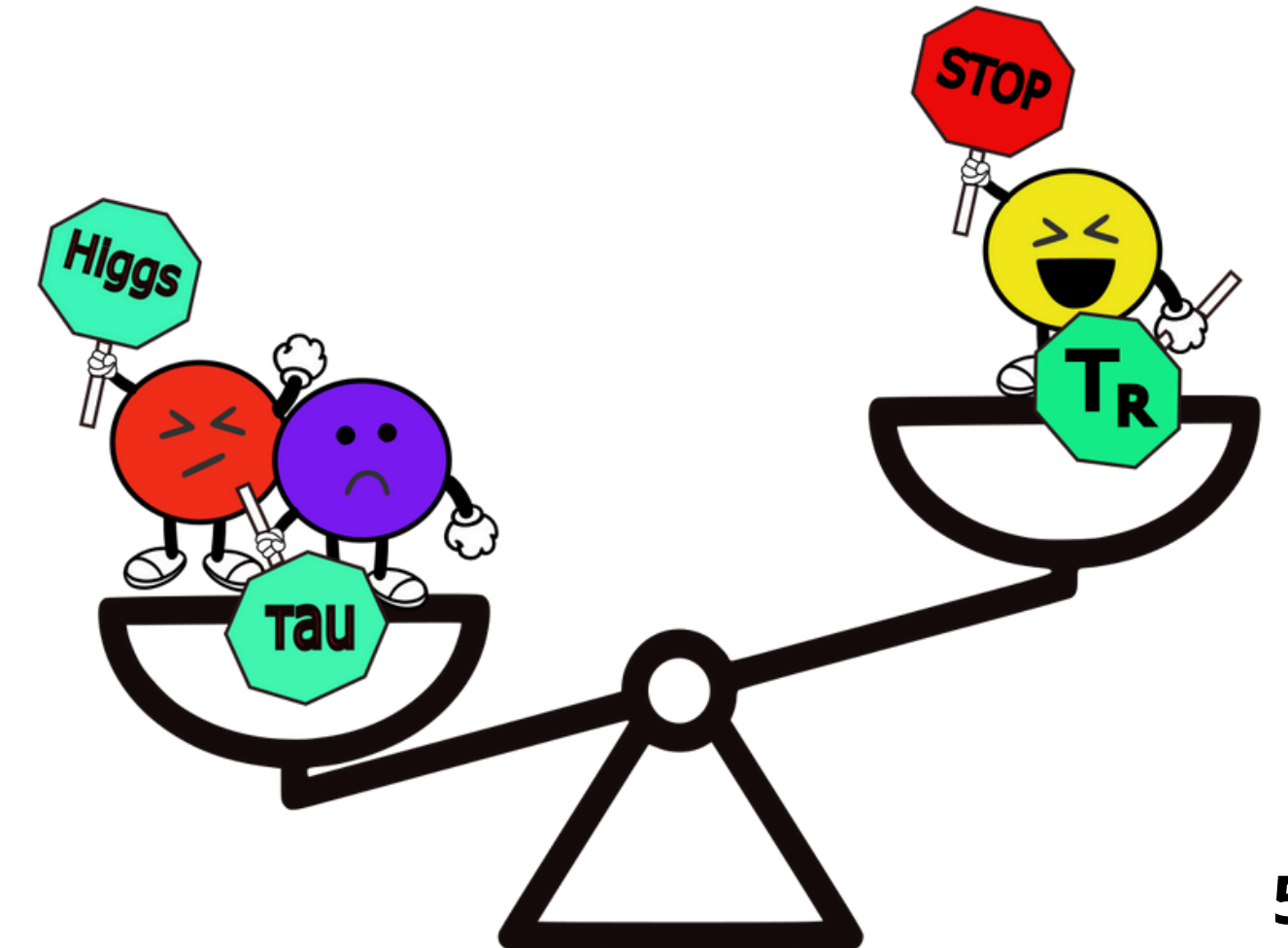
Thermally averaged cross section

Dilution effect of Universe expansion

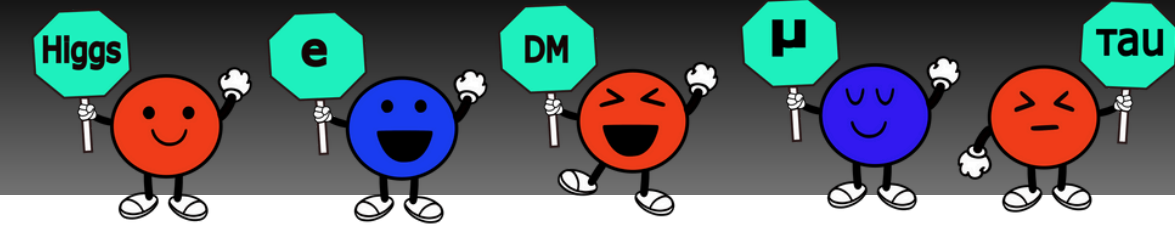
- Most important channels:



- Processes with \mathcal{T} and Higgs are forbidden, since $m_{\tau,h} \gg T_R$.



Boltzmann Equation



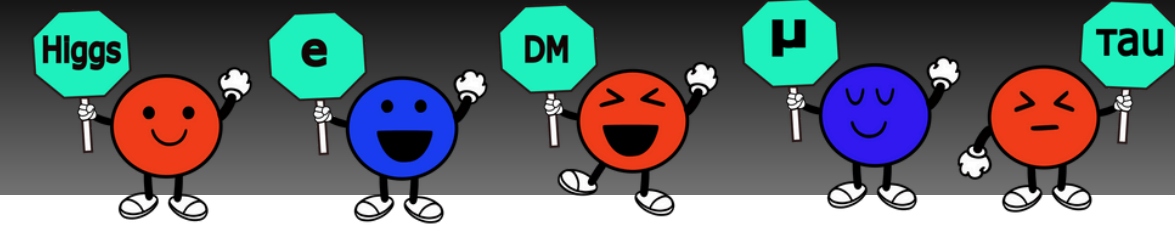
- The production cross section is given by

$$\sigma_{ff \rightarrow SS} = \frac{1}{2} \times \frac{1}{4} \times \frac{g^2 \lambda^2 m_f^2}{32\pi m_W^2 m_h^4} \sqrt{1 - \frac{4m_f^2}{s}}$$

$\frac{1}{2}$ → identical particles in the final state

$\frac{1}{4}$ → initial state spin averaging

Boltzmann Equation



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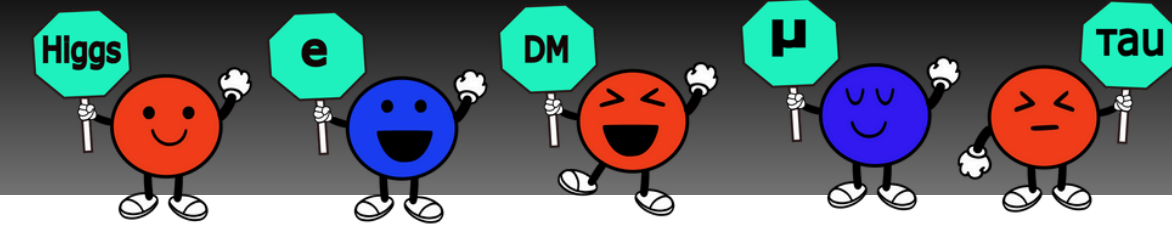
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- We can estimate the rate

$$\Gamma_e \approx \frac{g^2 \lambda^2 m_s^3 m_e^2 T^3}{256\pi^4 m_h^4 m_W^2} e^{-\frac{2m_s}{T}} \quad \Gamma_\mu \approx \frac{3g^2 \lambda^2 m_\mu^2 T^4}{512\pi^4 m_h^4 m_W^2} e^{-\frac{2m_\mu}{T}}$$

→ identical particles in the final state
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Boltzmann Equation



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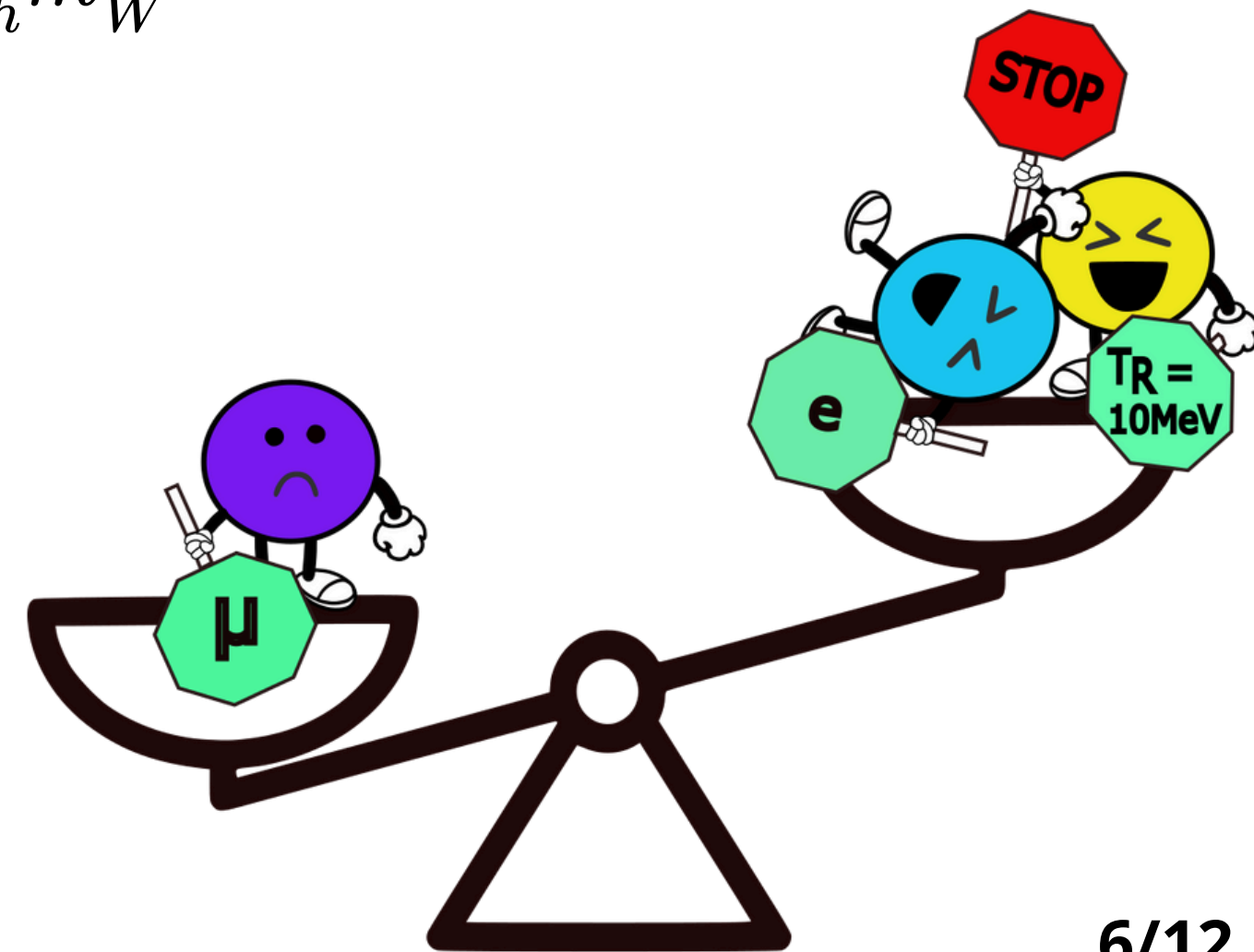
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- The ratio of the muon and electron rates for $T = T_R$ is

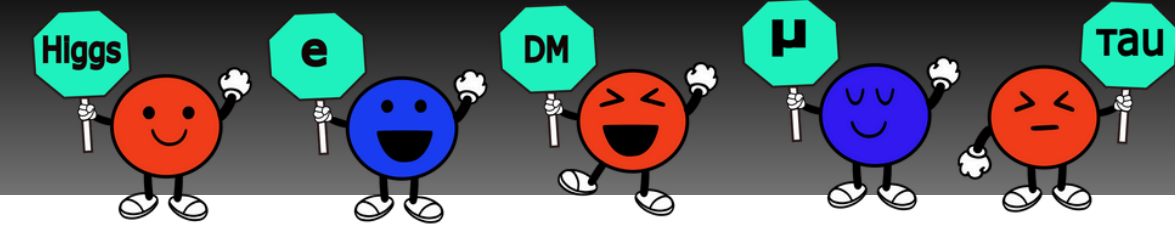
$$\frac{\Gamma_\mu}{\Gamma_e} \approx \frac{3}{2} \frac{m_\mu^4 T_R}{m_e^2 m_s^3} e^{-2(m_\mu - m_s)/T_R}$$

for $m_s = 50 MeV$ we obtain that Γ_μ dominate the contribution for

$$T_R > 10 MeV$$



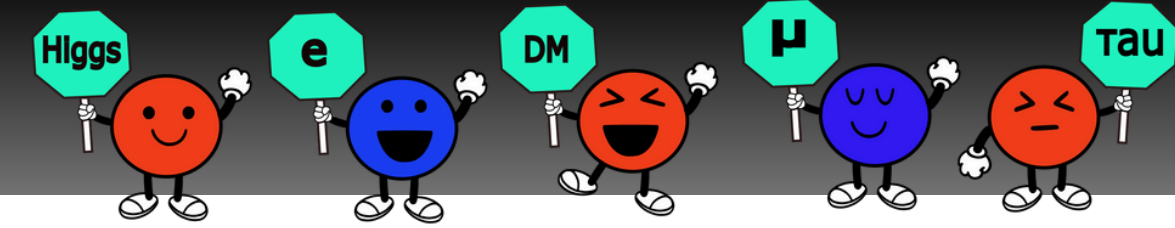
Boltzmann Equation



- For the *pure-freezing* scenario $\Gamma(SS \rightarrow ff) = 0$

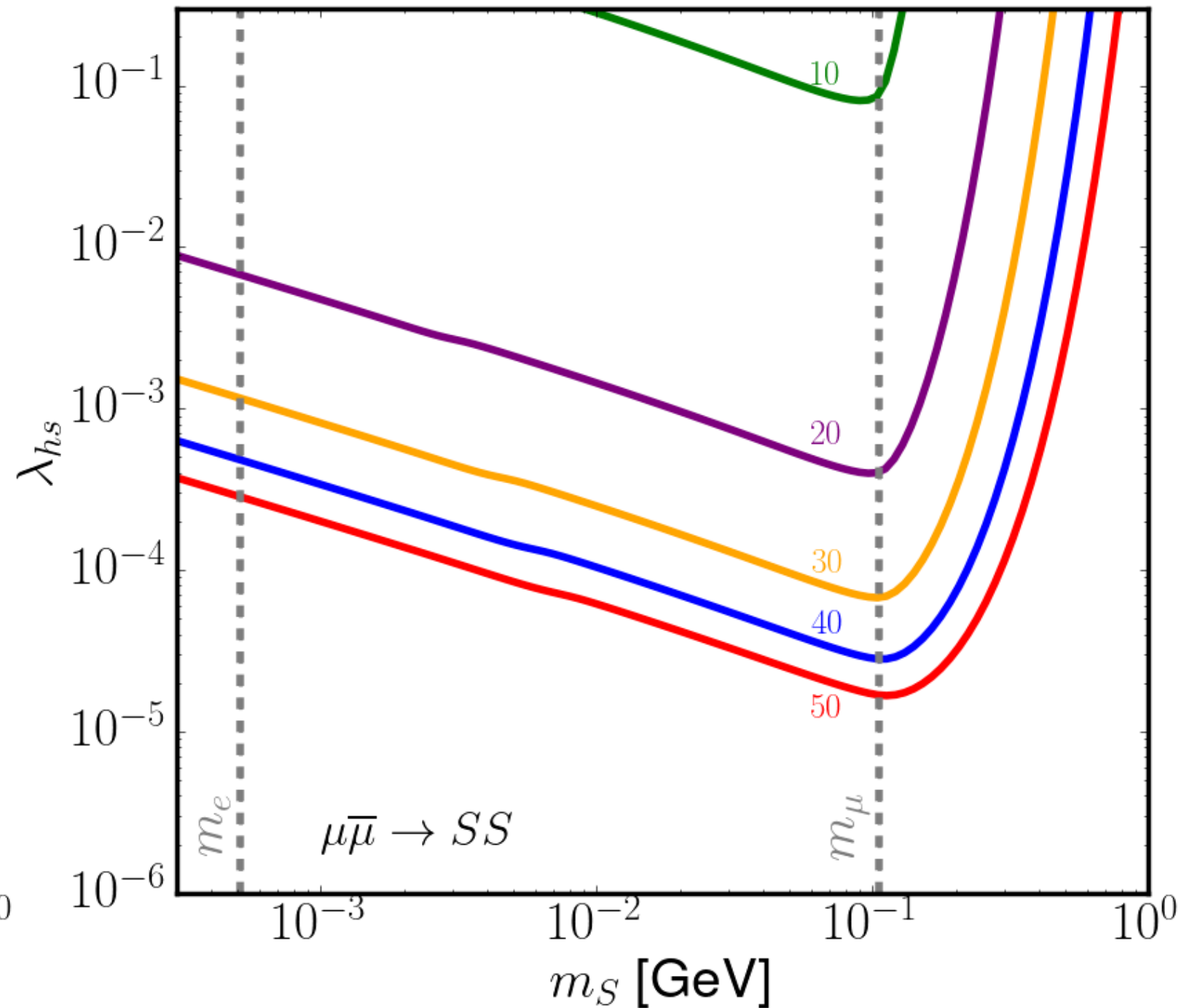
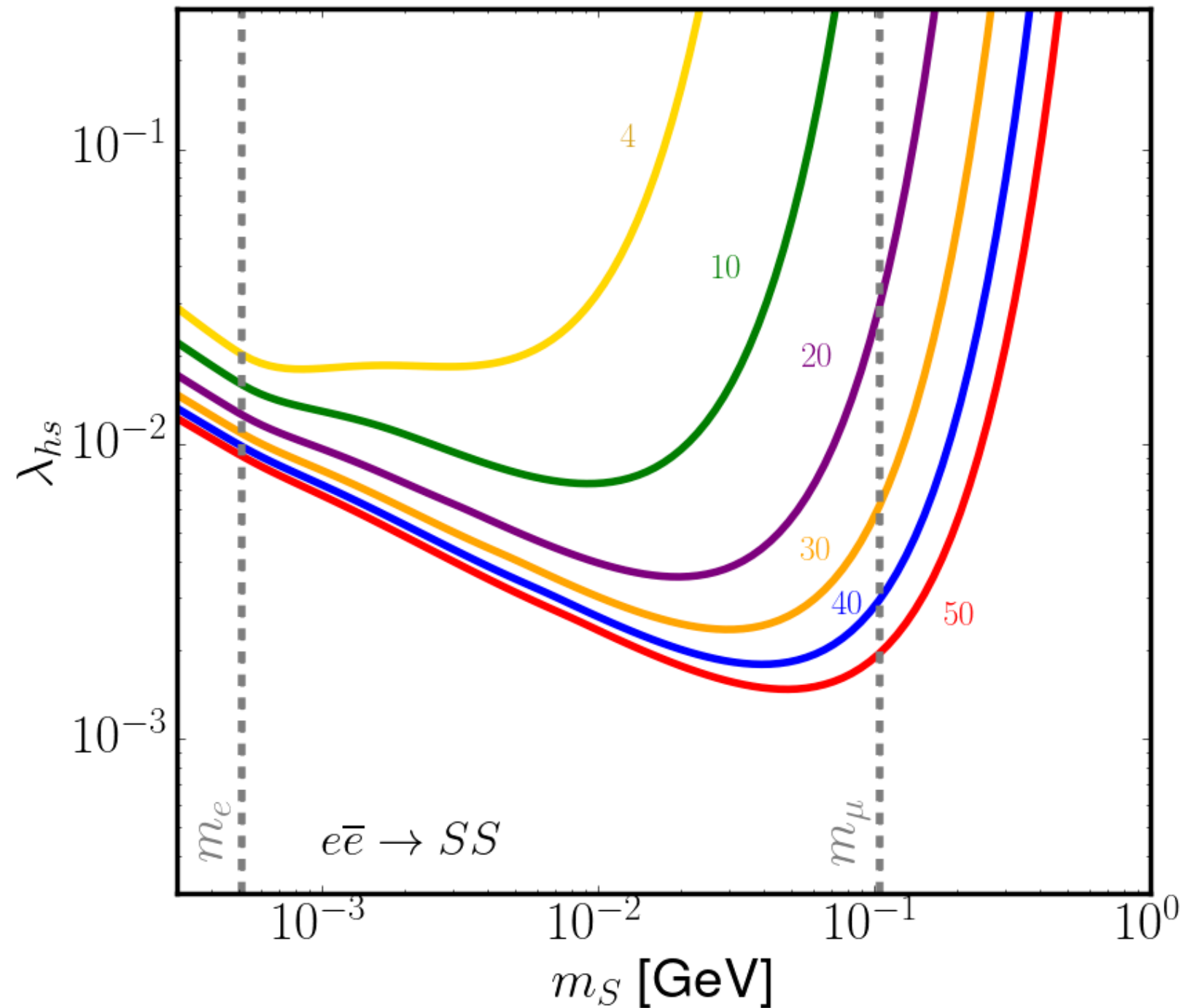
$$\frac{dY_S}{dx} = 2 \sqrt{\frac{8\pi^2 M_{Pl}^2}{45}} \frac{g_*^{1/2} m_s}{x^2} \sum_{f=e,\mu} \langle \sigma_{ff \rightarrow SS\nu} \rangle Y_f^{(eq)2}$$

Boltzmann Equation



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Constraints



- Invisible Higgs decay

$$BR_{inv} = \frac{\Gamma_{h \rightarrow ss}}{\Gamma_{SM}^{Tot} + \Gamma_{h \rightarrow ss}}$$

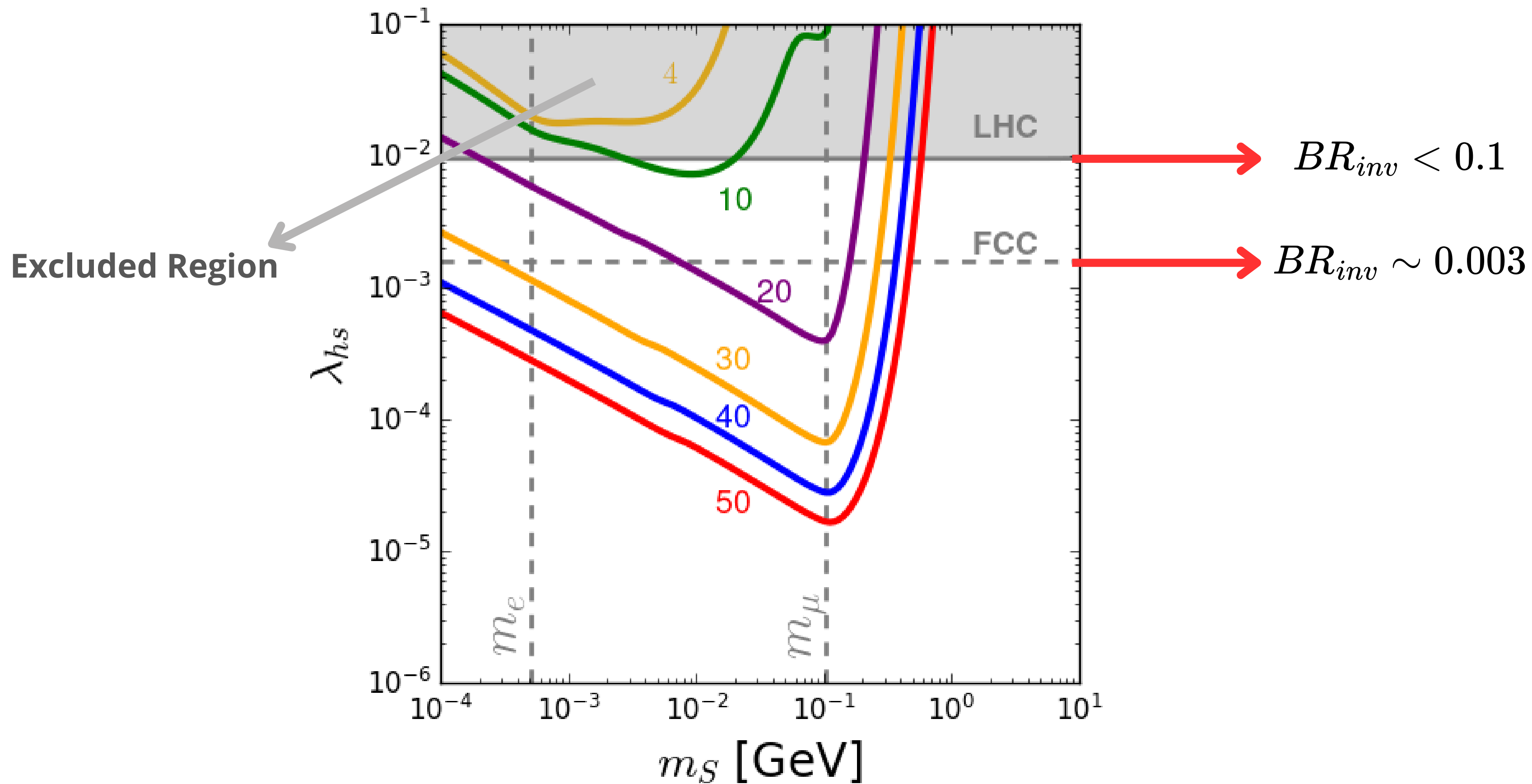
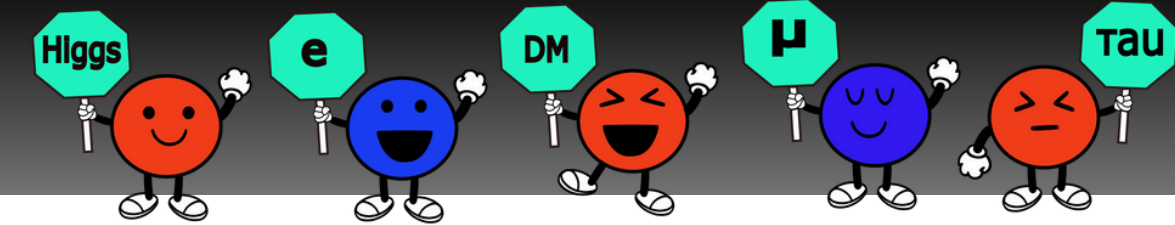
where

$$\Gamma_{h \rightarrow ss} = \frac{\lambda^2}{32\pi m_h} \sqrt{1 - \frac{4m_s^2}{m_h^2}}$$

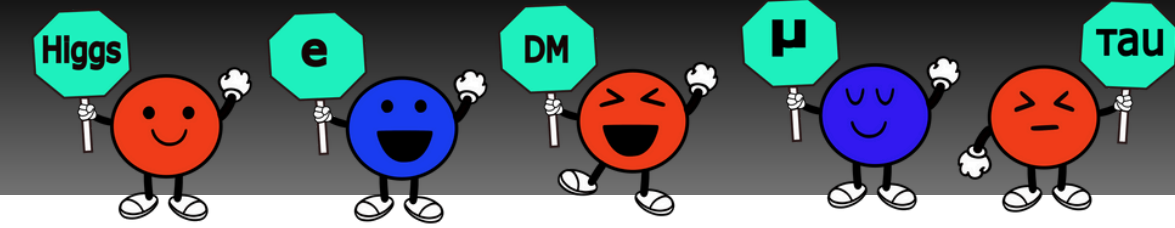
- We impose $BR_{inv} < 0.1$.
- The future collider FCC tuned as a Higgs factory can reach the limit of

$$BR_{inv} \sim 0.003$$

Results



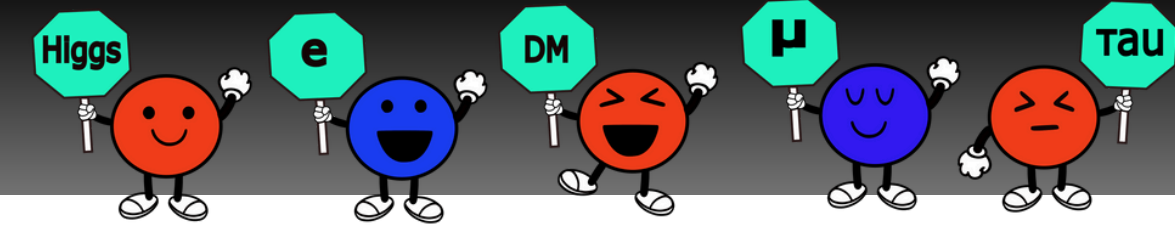
In progress...



- We will also consider the case of fermionic DM

$$L \supset \frac{gm_f}{2m_W} h \bar{f} f + \frac{\lambda_{hs} v_{SM}}{2} h \chi \bar{\chi}, \quad L \supset \frac{gm_f}{2m_W} h \bar{f} f + \frac{\lambda_{hs} v_{SM}}{2} i h \chi \gamma^5 \bar{\chi}$$

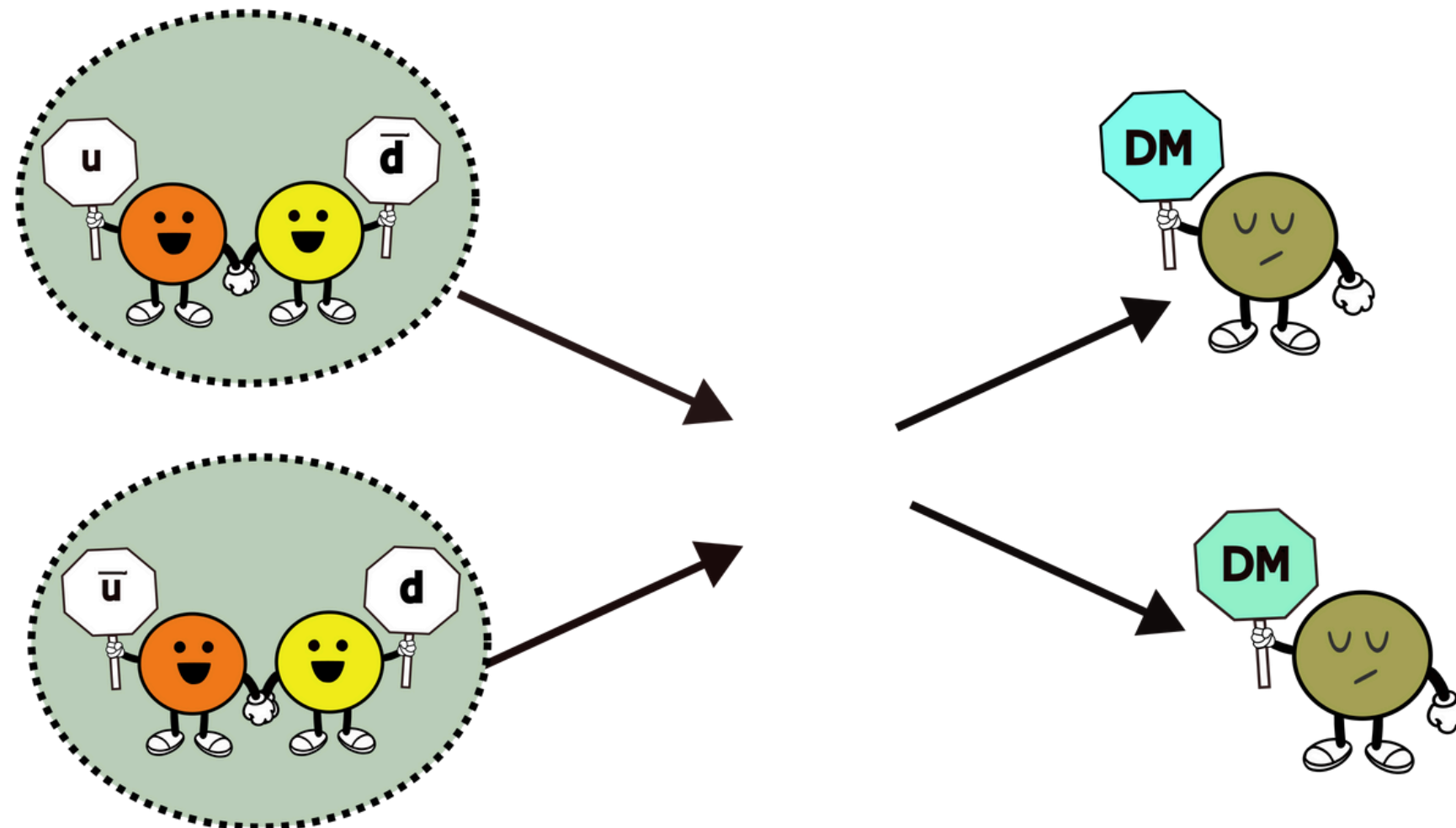
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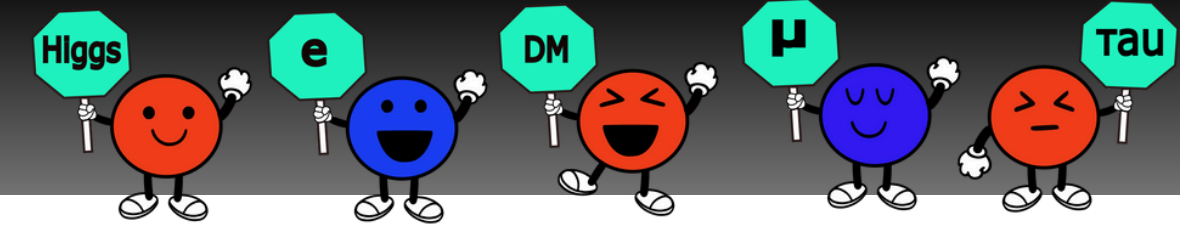


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$$L \supset \frac{gm_f}{2m_W} h \bar{f} f + \frac{\lambda_{hs} v_{SM}}{2} h \chi \bar{\chi}, \quad L \supset \frac{gm_f}{2m_W} h \bar{f} f + \frac{\lambda_{hs} v_{SM}}{2} i h \chi \gamma^5 \bar{\chi}$$

- We will also consider the mesons contributions





- **Freeze-in Mechanism:**
 - Dark Matter is Boltzmann suppressed if its mass exceeds the reheating temperature.
 - Correct DM abundance achieved with lower mass at low reheating temperatures.
- **Viability and Implications:**
 - Scenarios with low reheating temperatures are viable.
- **Future Implications**
 - Significant contribution from mesons after the Quantum Chromodynamics (QCD) phase transition.

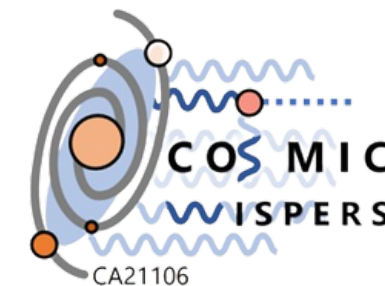
THANK YOU!



SCAN ME



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